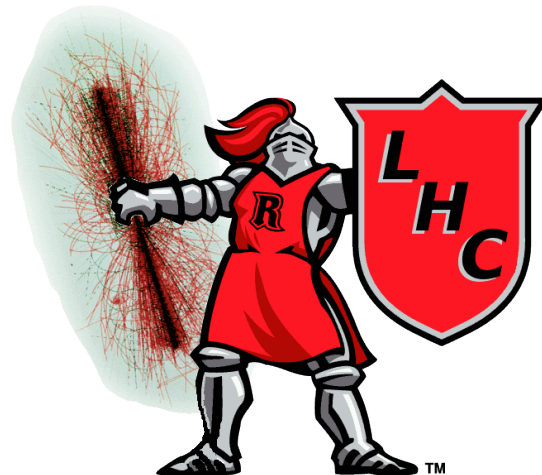


Searching for the Unexpected with Unsupervised Machine Learning

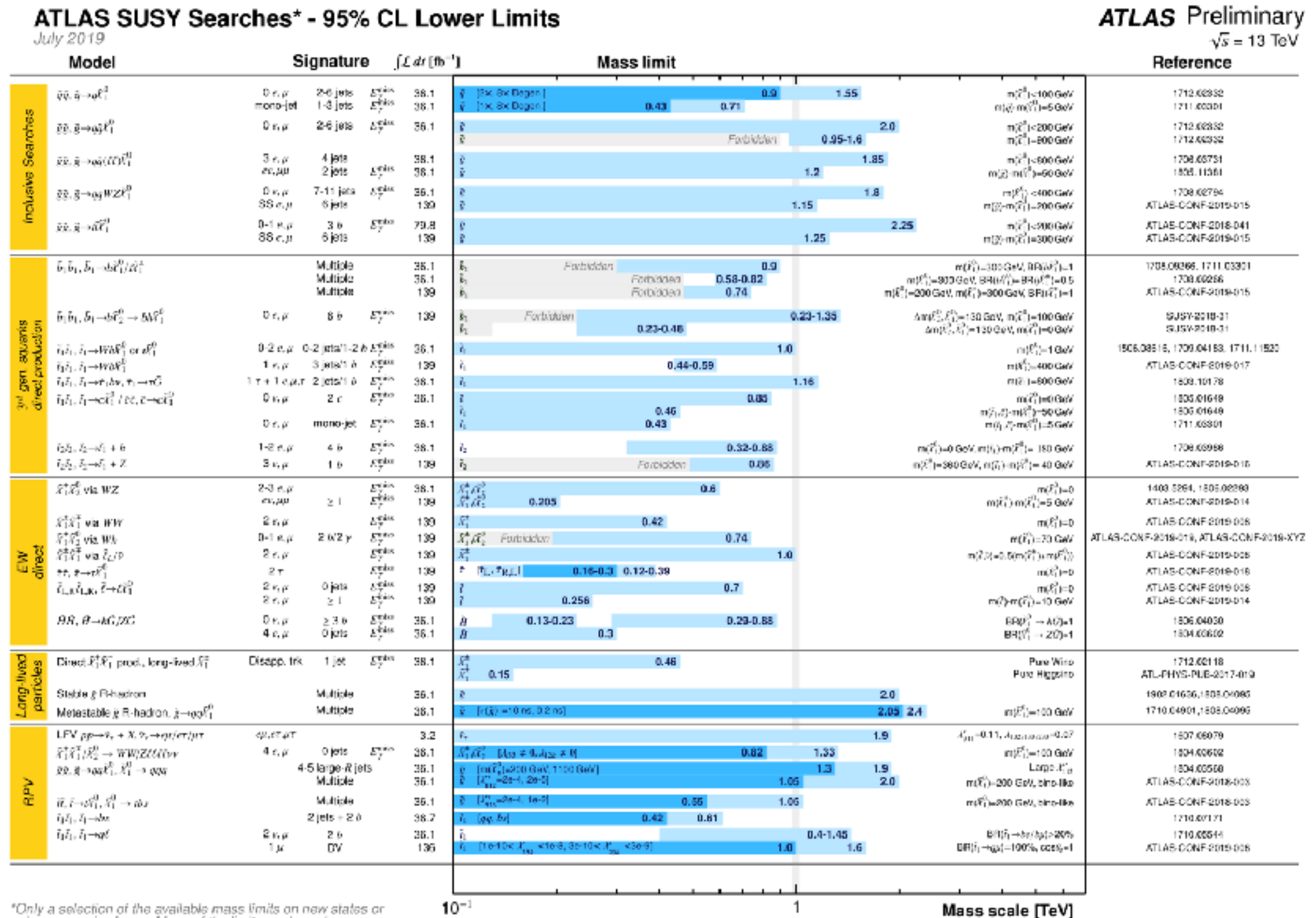
David Shih

Bay Area Particle Theory Seminar

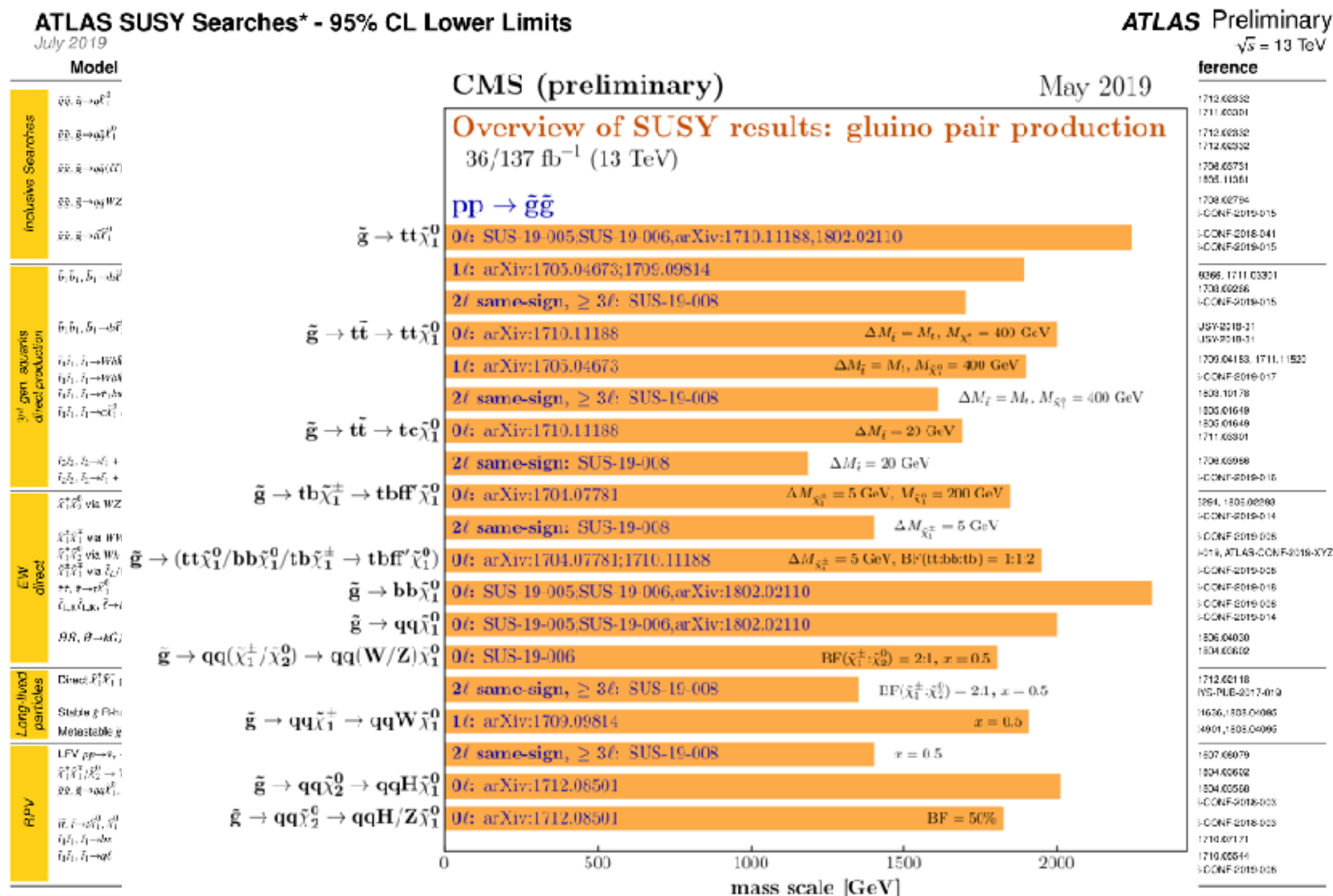
October 25, 2019



Current Status of NP Searches @ LHC



Current Status of NP Searches @ LHC



*Only a selection of phenomena is illustrated.

Selection of observed limits at 95% C.L. (theory uncertainties are not included). Probe **up to** the quoted mass limit for light LSPs unless stated otherwise. The quantities ΔM and x represent the absolute mass difference between the primary sparticle and the LSP, and the difference between the intermediate

Current Status of NP Searches @ LHC

ATLAS SUSY Searches* - 95% CL Lower Limits

July 2019

Model

$\tilde{q}\tilde{q}, \tilde{g} \rightarrow q\bar{q}$
 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}l\bar{l}$

CMS (preliminary)

Overview of SUSY results: gluino pair production

May 2019

ATLAS Preliminary

$\sqrt{s} = 13$ TeV

Reference

1712.02832
1711.02031
1712.02832
1712.02832

ATLAS Exotics Searches* - 95% CL Upper Exclusion Limits

Status: May 2019

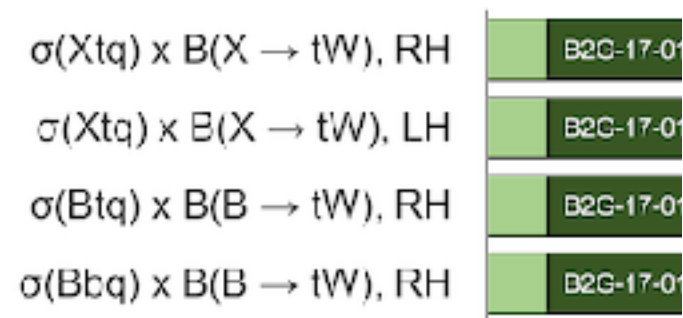
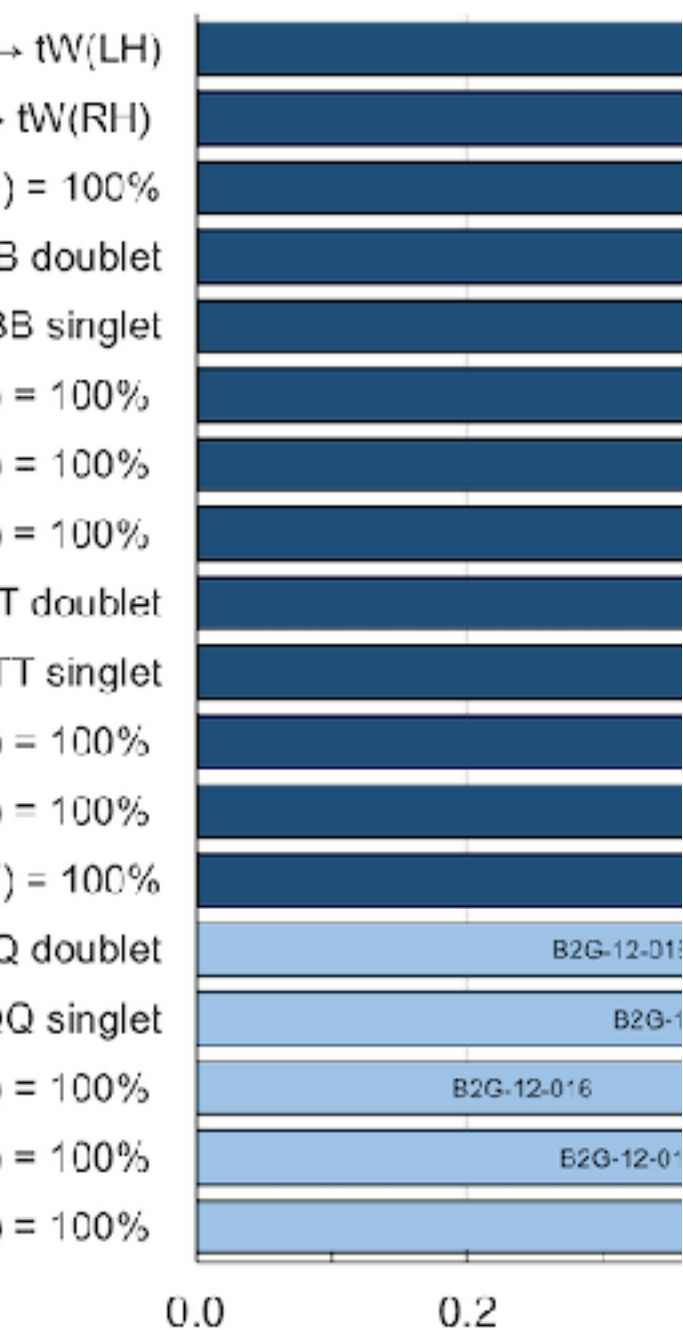
ATLAS Preliminary

$\sqrt{s} = 8, 13$ TeV

$\int \mathcal{L} dt = (3.2 - 139) \text{ fb}^{-1}$

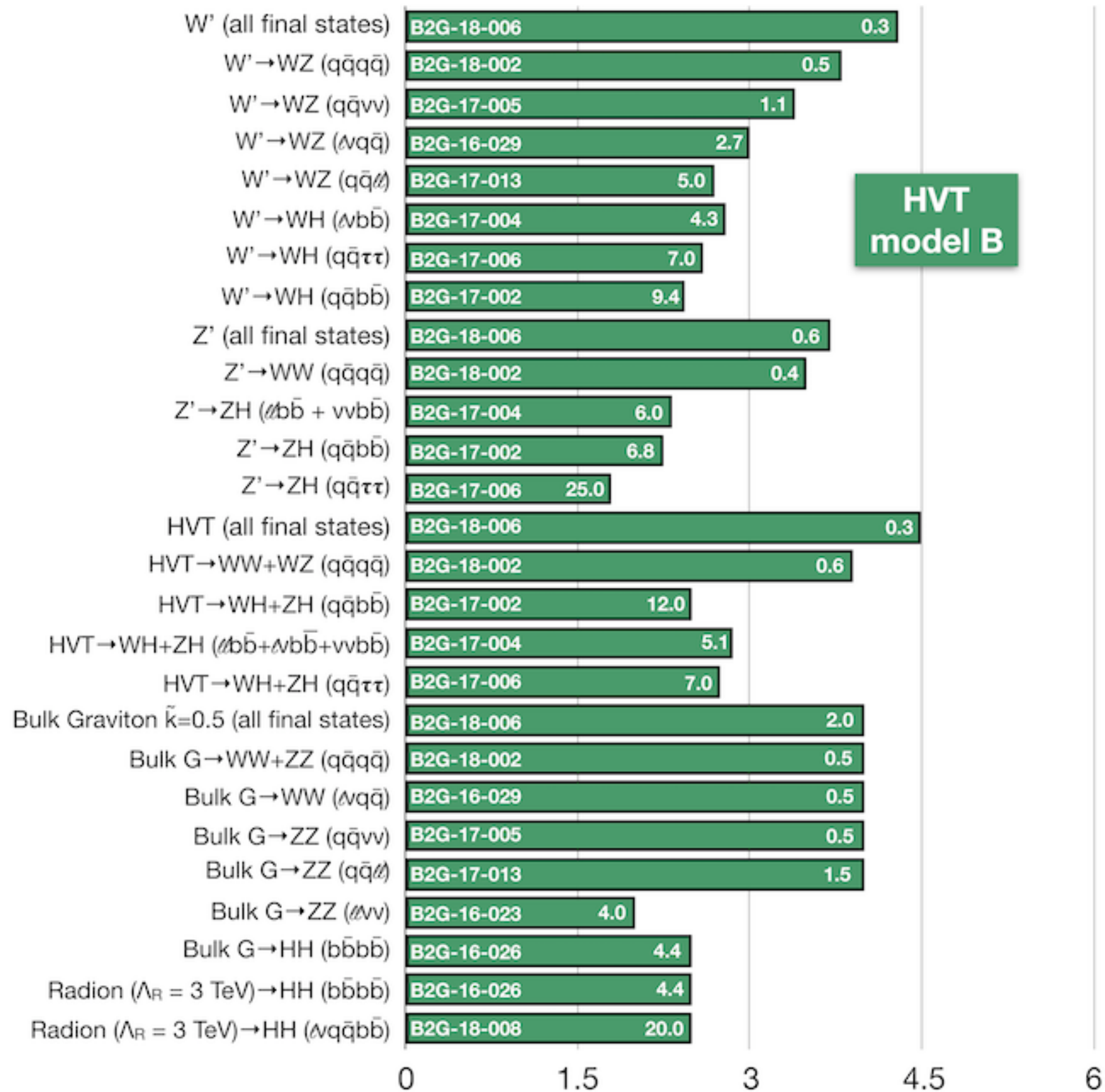
Model	ℓ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Limit	Reference		
Extra dimensions	ADD $G_{KK} \rightarrow g/q$	$0 \leq p_T$	$1-4$	Yes	35.1	M_{KK}	2.7 TeV	
	ADD non-resonant π_T	2	-	-	35.7	M_{KK}	0.6 TeV	
	ADD OBH	-	2	-	37.0	M_{KK}	8.9 TeV	
	ADD BH high $\sum p_T$	$> 1 \leq p_T$	> 2	-	3.2	M_{KK}	0.2 TeV	
	ADD BH multijet	-	≥ 3	-	3.6	M_{KK}	9.55 TeV	
	RS1 $G_{KK} \rightarrow \pi_T$	2	-	-	35.7	$G_{KK} \text{ mass}$	4.1 TeV	
	Bulk RS $G_{KK} \rightarrow WW/ZZ$	multi-channel	-	-	35.1	$G_{KK} \text{ mass}$	2.3 TeV	
	Bulk RS $G_{KK} \rightarrow WW \rightarrow q\bar{q}q\bar{q}$	$0 \leq p_T$	2	-	189	$G_{KK} \text{ mass}$	1.6 TeV	
	Bulk RS $G_{KK} \rightarrow \pi_T$	$1 \leq p_T$	> 1	> 1.5	Yes	35.1	$G_{KK} \text{ mass}$	3.8 TeV
	ZUED + RPP	$1 \leq p_T$	≥ 2	≥ 3	Yes	35.1	$G_{KK} \text{ mass}$	1.8 TeV
Gauge bosons	SSM $Z' \rightarrow \ell\bar{\ell}$	$2 \leq p_T$	-	-	189	$Z' \text{ mass}$	5.1 TeV	
	SSM $Z' \rightarrow \nu\bar{\nu}$	2	-	-	35.1	$Z' \text{ mass}$	2.42 TeV	
	Leptophobic $Z' \rightarrow b\bar{b}$	-	2	-	35.1	$Z' \text{ mass}$	2.1 TeV	
	Leptophobic $Z' \rightarrow t\bar{t}$	$1 \leq p_T$	> 1	> 1.5	Yes	35.1	$Z' \text{ mass}$	3.0 TeV
	SSM $W' \rightarrow \ell\nu$	$1 \leq p_T$	-	Yes	189	$W' \text{ mass}$	6.0 TeV	
	SSM $W' \rightarrow \nu\nu$	1	-	Yes	35.1	$W' \text{ mass}$	3.7 TeV	
	HVT $V' \rightarrow WZ \rightarrow q\bar{q}q\bar{q}$ model B	$0 \leq p_T$	2	-	189	$V' \text{ mass}$	3.6 TeV	
	HVT $V' \rightarrow WH/ZH$ model B	multi-channel	-	-	35.1	$V' \text{ mass}$	2.90 TeV	
	LRSM $W_R \rightarrow \ell\nu$	multi-channel	-	-	35.1	$W_R \text{ mass}$	3.25 TeV	
	LRSM $W_R \rightarrow \mu\nu$	2	-	-	80	$W_R \text{ mass}$	5.0 TeV	
CT	Cl qqg	-	2	-	37.0	A	21.8 TeV	
	Cl Wqq	$2 \leq p_T$	-	-	35.1	A	40.0 TeV	
	Cl ttg	$\geq 1 \leq p_T$	≥ 1	≥ 1	Yes	35.1	A	2.57 TeV
DM	Adax vector mediator (Dirac DM)	$0 \leq p_T$	$1-4$	Yes	35.1	m_{Adax}	1.55 TeV	
	Colored scalar mediator (Dirac DM)	$0 \leq p_T$	$1-4$	Yes	35.1	m_{Adax}	1.57 TeV	
	WV(L) LFI (Dirac DM)	$0 \leq p_T$	1	≤ 1	Yes	35.1	m_L	700 GeV
	Scalar reson. $\phi \rightarrow \ell\bar{\ell}$ (Dirac DM)	$0-1 \leq p_T$	1	< 1	Yes	35.1	m_ϕ	3.4 TeV
LO	Scalar LQ 1 st gen	$1, 2 \leq p_T$	≥ 2	Yes	35.1	$LQ \text{ mass}$	1.4 TeV	
	Scalar LQ 2 nd gen	$1, 2 \leq p_T$	≥ 2	Yes	35.1	$LQ \text{ mass}$	1.55 TeV	
	Scalar LQ 3 rd gen	2	2	-	35.1	$LQ \text{ mass}$	1.03 TeV	
	Scalar LQ 3 rd gen	$0-1 \leq p_T$	2	Yes	35.1	$LQ \text{ mass}$	970 GeV	
Heavy quarks	VLO $11 \rightarrow 11/Z/\gamma Wb + X$	multi-channel	-	-	35.1	$\Gamma \text{ mass}$	1.37 TeV	
	VLO $66 \rightarrow 66/Zb + X$	multi-channel	-	-	35.1	$\Gamma \text{ mass}$	1.34 TeV	
	VLO $T_{KK} T_{KK} \rightarrow Wb + X$	$2(SS)/3(S) \leq p_T$	≥ 1	≥ 1	Yes	35.1	$T_{KK} \text{ mass}$	1.54 TeV
	VLO $Y \rightarrow Wb + X$	$1 \leq p_T$	> 1	> 1	Yes	35.1	$Y \text{ mass}$	1.85 TeV
	VLO $B \rightarrow Wb + X$	$0 \leq p_T$	2	≥ 1	Yes	79.8	$B \text{ mass}$	1.21 TeV
	VLO $QQ \rightarrow WbWq$	$1 \leq p_T$	> 4	Yes	20.3	$Q \text{ mass}$	690 GeV	
Exotic fermions	Exotic quark $q' \rightarrow qg$	-	2	-	189	$q' \text{ mass}$	6.7 TeV	
	Exotic quark $q' \rightarrow q\gamma$	1	1	-	35.7	$q' \text{ mass}$	5.3 TeV	
	Exotic quark $b' \rightarrow Ag$	-	1	> 1	Yes	35.1	$b' \text{ mass}$	2.6 TeV
	Exotic lepton ℓ'	$3 \leq p_T$	-	-	20.3	$\ell' \text{ mass}$	3.0 TeV	
	Exotic lepton ν'	$3 \leq p_T$	-	-	20.3	$\nu' \text{ mass}$	1.6 TeV	
Other	Type III Seesaw	$1 \leq p_T$	≥ 2	Yes	79.8	$N^0 \text{ mass}$	560 GeV	
	LRSM Majorana ν	2	2	-	35.1	$N^0 \text{ mass}$	3.2 TeV	
	Higgs triplet $H^{++} \rightarrow \ell\bar{\ell}$	$2, 3, 4 \leq p_T$ (SS)	-	-	35.1	$H^{++} \text{ mass}$	870 GeV	
	Higgs triplet $H^{++} \rightarrow \ell\bar{\ell}$	$3 \leq p_T$	-	-	20.3	$H^{++} \text{ mass}$	400 GeV	
	Multi-charged particles	-	-	-	35.1	multi-charged particle mass	1.22 TeV	

Vector-like quark single production

 V_e 

HEP 2019

The quantities ΔM and r represent the mass of the sparticle and the LSP relative to ΔM , i.e.

$$\text{BB} = (l^{\pm}, l^{\pm} l^{\pm})$$
Resonances to dibosons ($\sqrt{s} = 13$ TeV)

CMS, EPS-HEP 2019

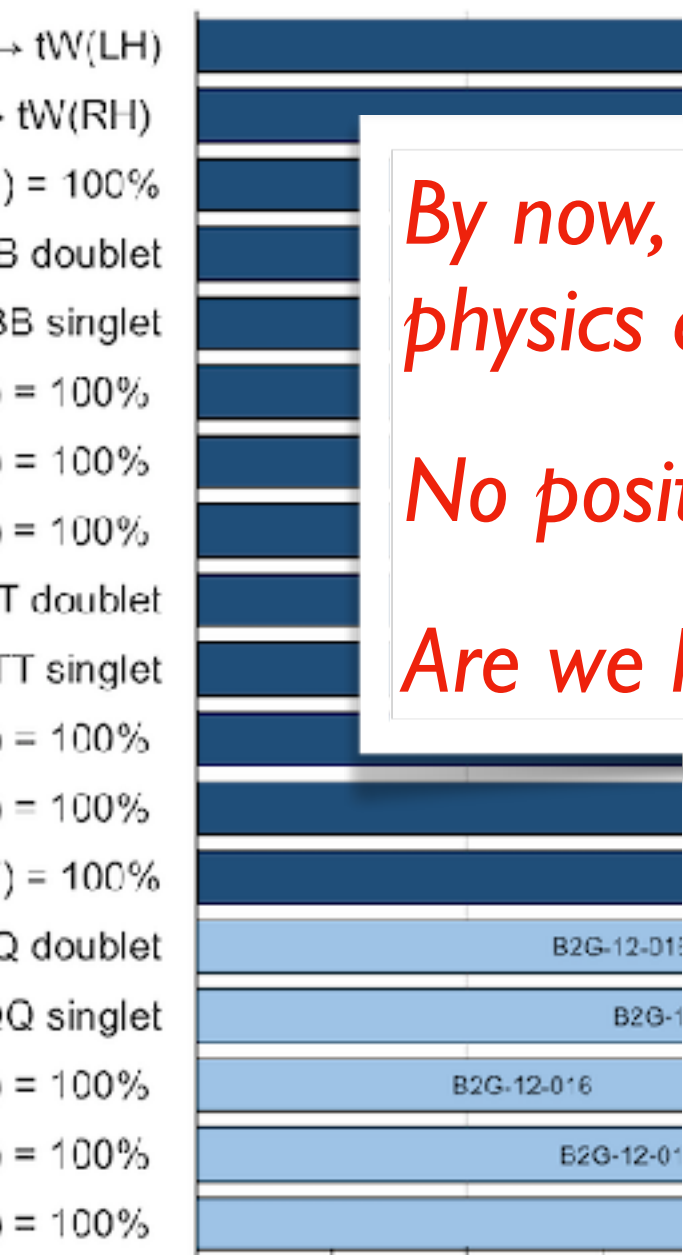
95% CL Lower Mass Limit [TeV]

(4) Upper Ocean Circulation Limit [ft.]

Vector-like quark single production

$\sigma(Xtq) \times B(X \rightarrow tW), RH$	B2G-17-01
$\sigma(Xtq) \times B(X \rightarrow tW), LH$	B2G-17-01
$\sigma(Btq) \times B(B \rightarrow tW), RH$	B2G-17-01
$\sigma(Bbq) \times B(B \rightarrow tW), RH$	B2G-17-01

Ve



Resonances to dibosons ($\sqrt{s} = 13 \text{ TeV}$)

W' (all final states)	B2G-18-006	0.3
$W' \rightarrow WZ (q\bar{q}q\bar{q})$	B2G-18-002	0.5
$W' \rightarrow WZ (q\bar{q}\nu\nu)$	B2G-17-005	1.1
$W' \rightarrow WZ (\nu q\bar{q})$	B2G-16-029	2.7
$W' \rightarrow WZ (q\bar{q}\ell\ell)$	B2G-17-013	5.0
$W' \rightarrow WH (\nu b\bar{b})$	B2G-17-004	4.3
$W' \rightarrow WH (q\bar{q}\tau\tau)$	B2G-17-006	7.0
$W' \rightarrow WH (q\bar{q}b\bar{b})$	B2G-17-002	9.4

HVT model B

By now, hundreds (thousands?) of searches for new physics at the LHC.

No positive signals so far, only limits.

Are we looking in the right places?

Bulk Graviton $k=0.5$ (all final states)	B2G-18-006	2.0
Bulk $G \rightarrow WW+ZZ (q\bar{q}q\bar{q})$	B2G-18-002	0.5
Bulk $G \rightarrow WW (\nu q\bar{q})$	B2G-16-029	0.5
Bulk $G \rightarrow ZZ (q\bar{q}\nu\nu)$	B2G-17-005	0.5
Bulk $G \rightarrow ZZ (q\bar{q}\ell\ell)$	B2G-17-013	1.5
Bulk $G \rightarrow ZZ (\ell\ell\nu\nu)$	B2G-16-023	4.0
Bulk $G \rightarrow HH (b\bar{b}b\bar{b})$	B2G-16-026	4.4
Radion ($\Lambda_R = 3 \text{ TeV}$) $\rightarrow HH (b\bar{b}b\bar{b})$	B2G-16-026	4.4
Radion ($\Lambda_R = 3 \text{ TeV}$) $\rightarrow HH (\nu q\bar{q}b\bar{b})$	B2G-18-008	20.0

CMS, EPS-HEP 2019

95% CL Lower Mass Limit [TeV]

HEP 2019

Selection of observed limits at 95% CL. The quantities ΔM and σ represent the mass difference and the cross-section relative to ΔM , or

$BB \rightarrow (l\bar{l}, l\bar{l})$
 $HH \rightarrow (l\bar{l}, l\bar{l})$

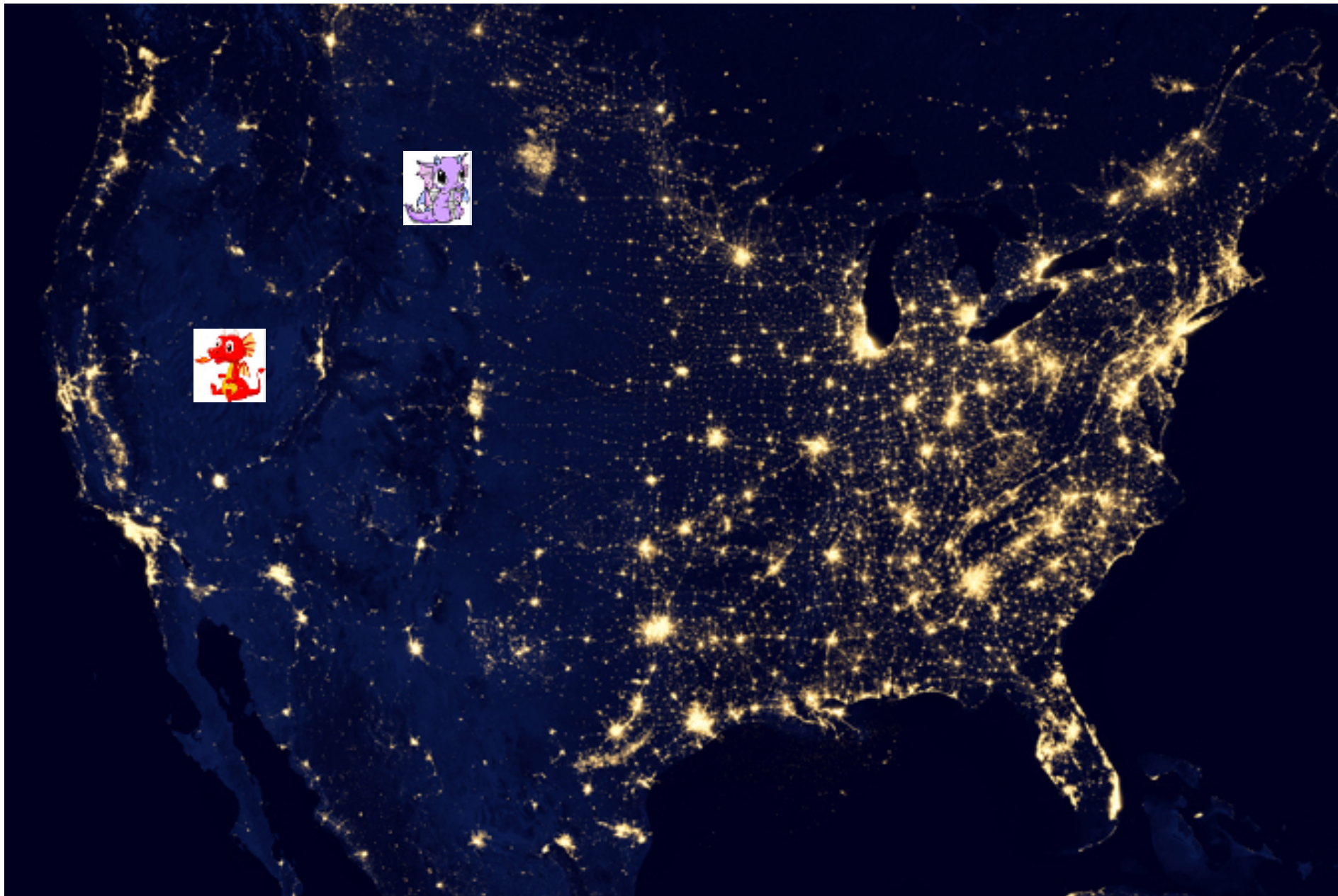
Our current coverage



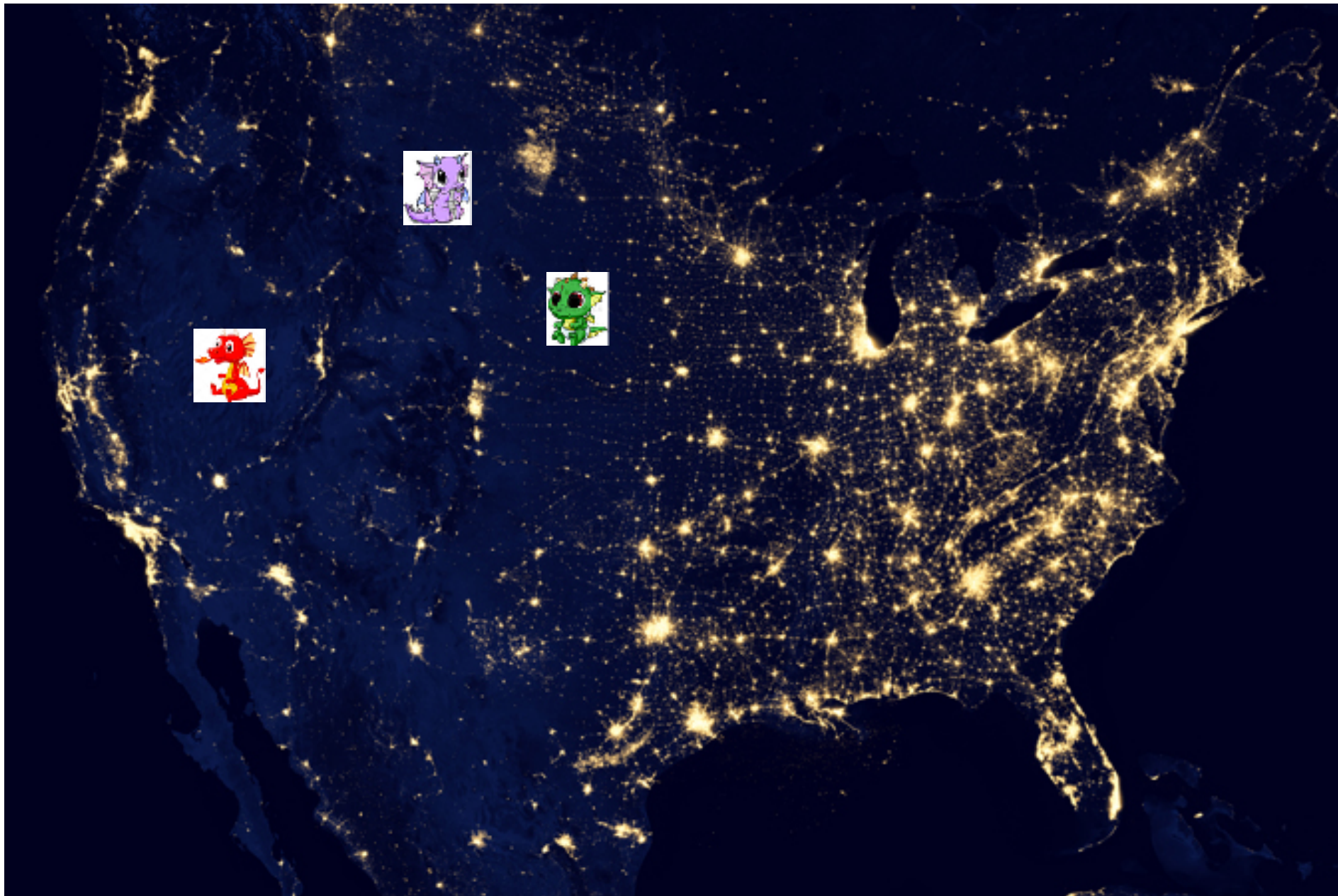
Our current coverage



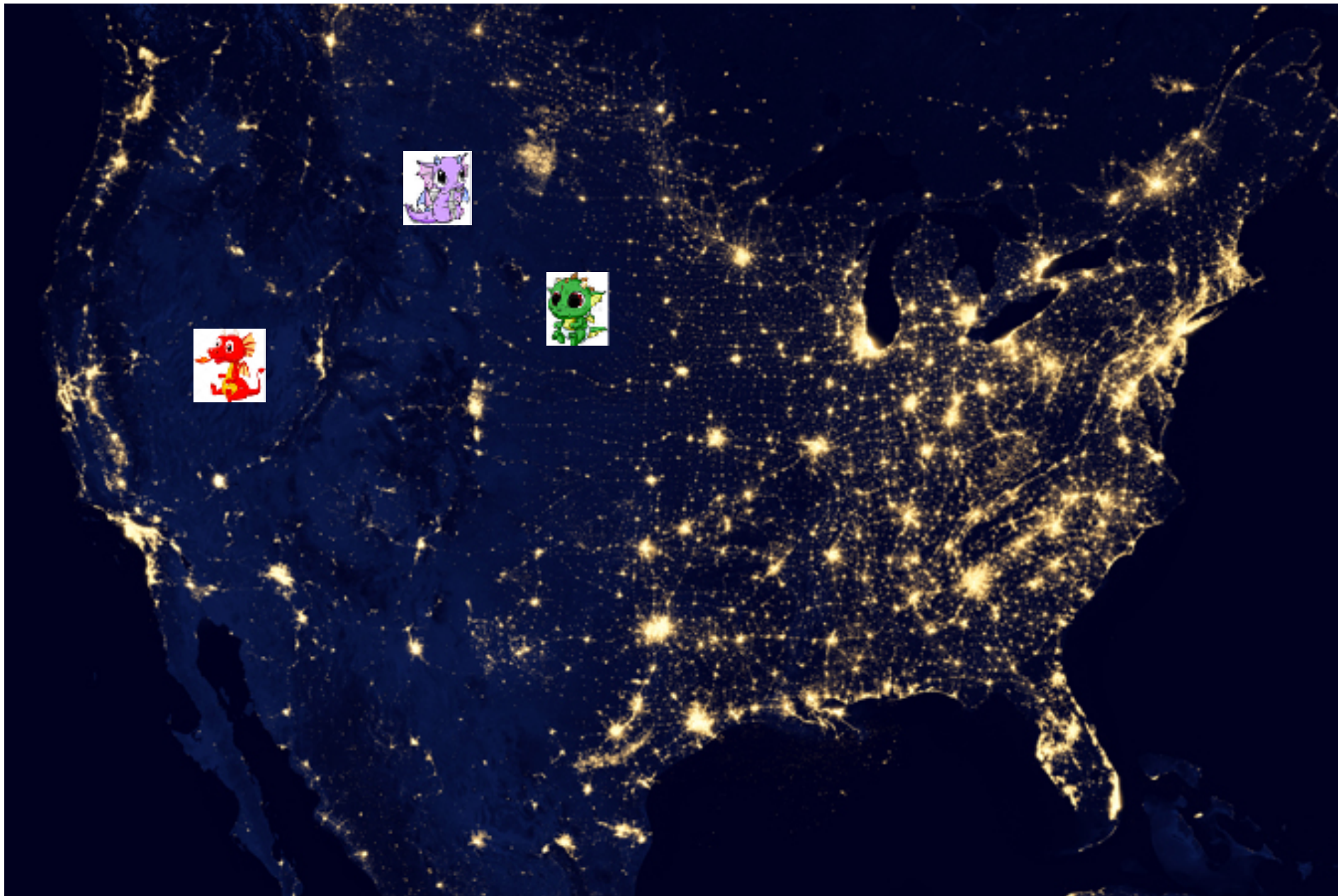
Our current coverage



Our current coverage



Our current coverage



How can we make sure we don't miss anything?

Model Independent Searches

Can we search for new physics without a model in mind?

Model Independent Searches

Can we search for new physics without a model in mind?

Yes?? A brief history of model independent searches in HEP:

- **D0** *“Sleuth”* PRD 62:092004 (2000)
PRD 64:012004 (2001)
PRL 86:3712 (2001)
- **H1 (Hera)** *“General Search”* PLB 602:14-30 (2004)
0705.3721
- **CDF** *“Sleuth/Vista”* 0712.1311 PRD 78:012002 (2008)
0712.2534 (submitted to PRL, NEVER PUBLISHED)
0809.3781 PRD 79:011101 (2009)
- **CMS** *“MUSIC”* CMS-PAS-EXO-14-016
- **ATLAS** *“Model independent general search”* 1807.07447 EPJC 79:120 (2019)



Model independent searches

The general approach behind all of these:

- Bin the data into exclusive final states,
- Compare a zillion 1D histograms between data and simulation (or just look at high p_T tails),
- Look for discrepancies.

From CDF 0712.2534:

A global comparison of data to standard model prediction is made in 16,486 kinematic distributions in 344 populated exclusive final states. In each final state, the

bottom quark (b), and missing momentum ($\cancel{p}$). Monte Carlo event generators are used to determine the standard model prediction. VISTA partitions data and Monte

Major drawbacks

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Over-reliance on simulation for background prediction

An example of what is found

CDF Run II (927 pb⁻¹)

Sleuth@CDFII result

(top 5)

fraction of pseudo experiments in this
final state as interesting as CDF data

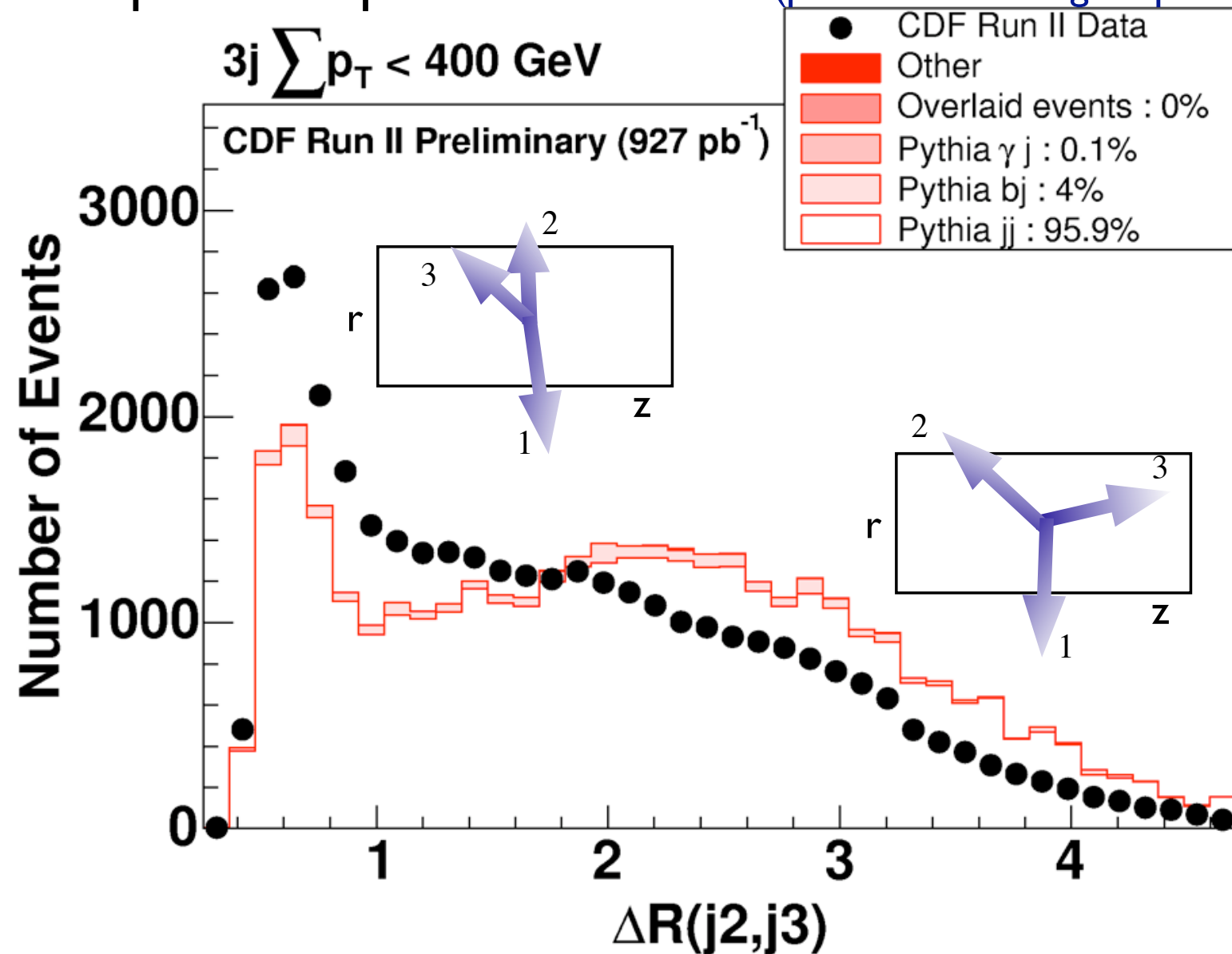
fraction of pseudo experiments in any
final state as interesting as CDF data

SLEUTH Final State	$\mathcal{P}$	
$b\bar{b}$	0.0055	$\tilde{\mathcal{P}} = 0.46$ 46% of pseudo experiments are expected to be as interesting Sleuth finds no significant excess in CDF Run II high-p _T data This does not prove there is no new physics present
$j\cancel{p}$	0.0092	
$\ell^+\ell'^+\cancel{p}jj$	0.011	
$\ell^+\ell'^+\cancel{p}$	0.016	
$\tau\cancel{p}$	0.016	

From B. Knuteson talk at UMich (2008)

An example of what is found

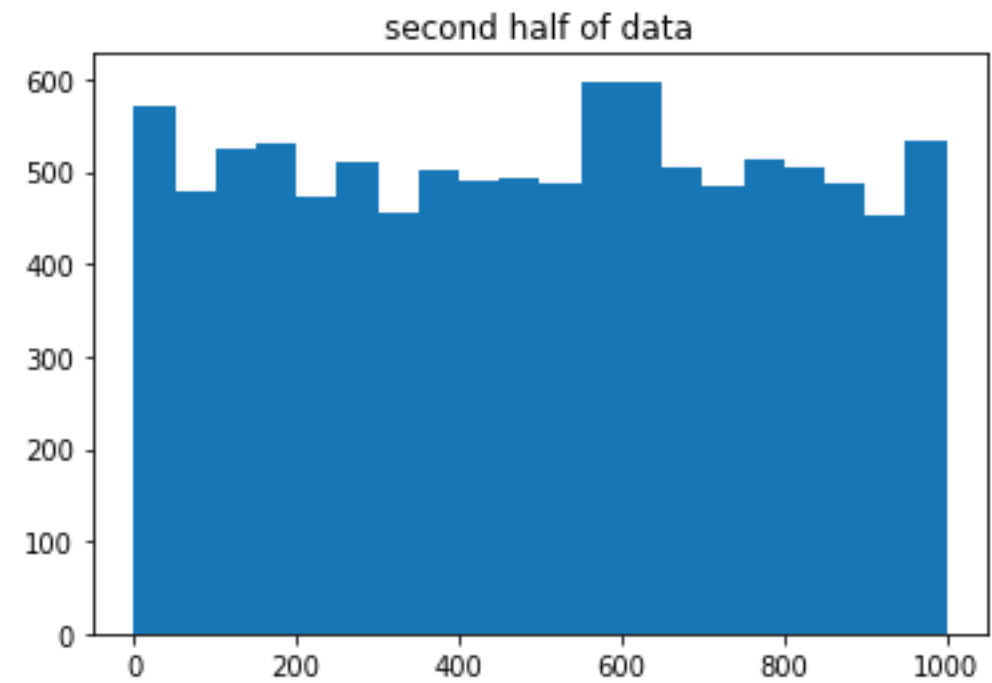
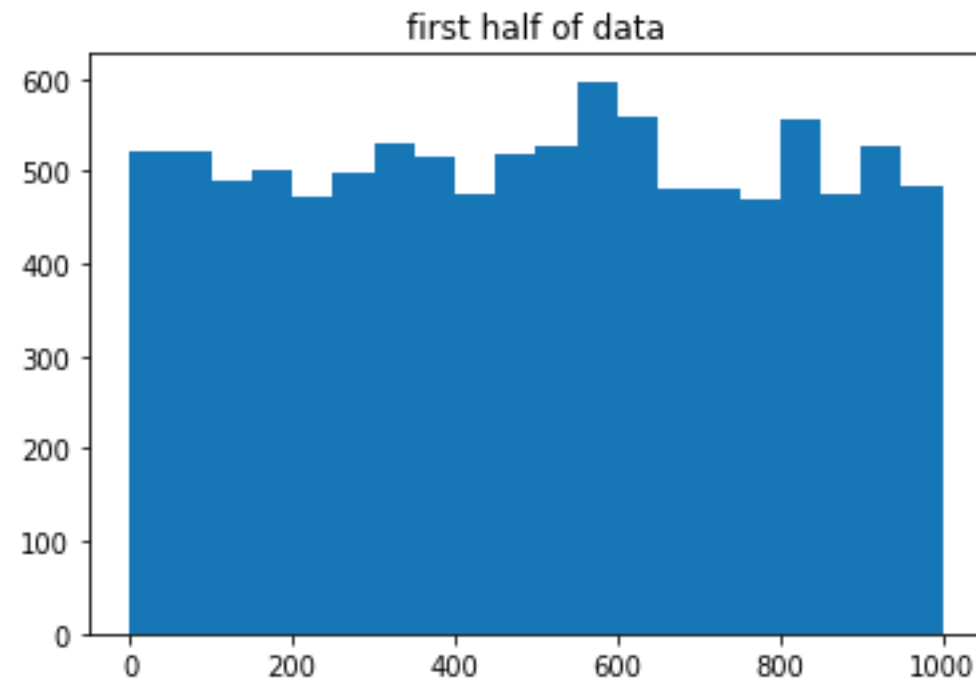
Sample discrepant distribution (parton showering suspected)



From B. Knuteson talk at UMich (2008)

Look elsewhere effect

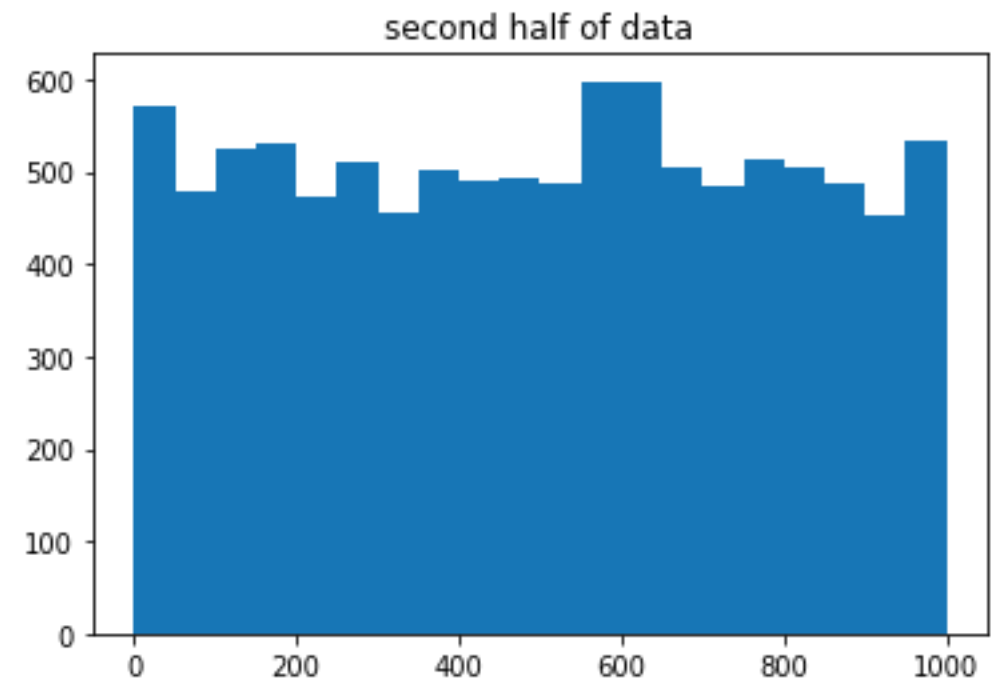
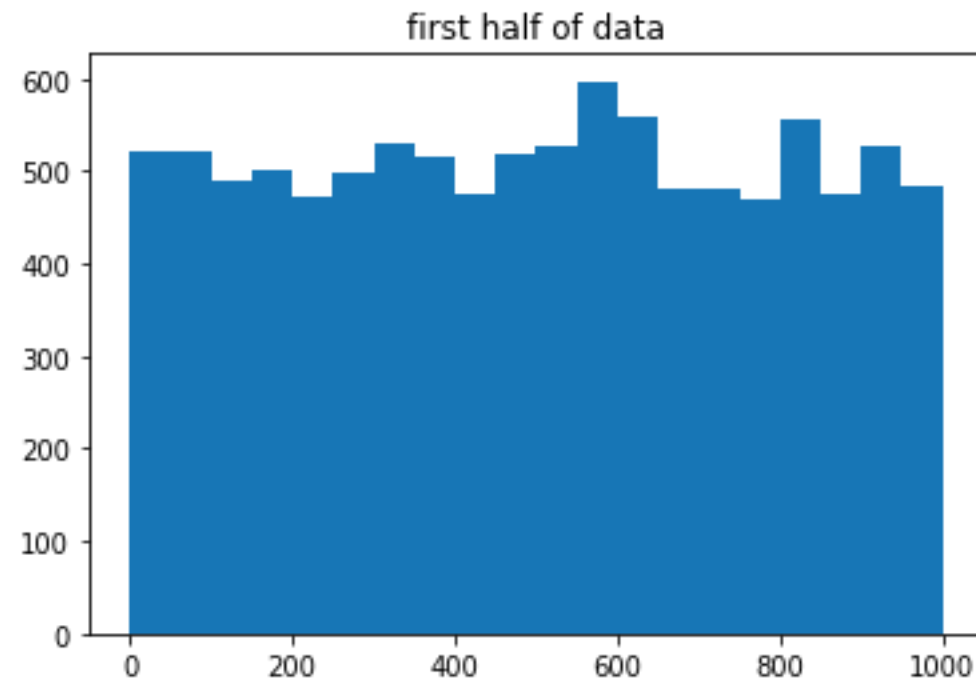
Look elsewhere effect



Simple way to mitigate LEE:

- Divide data in two

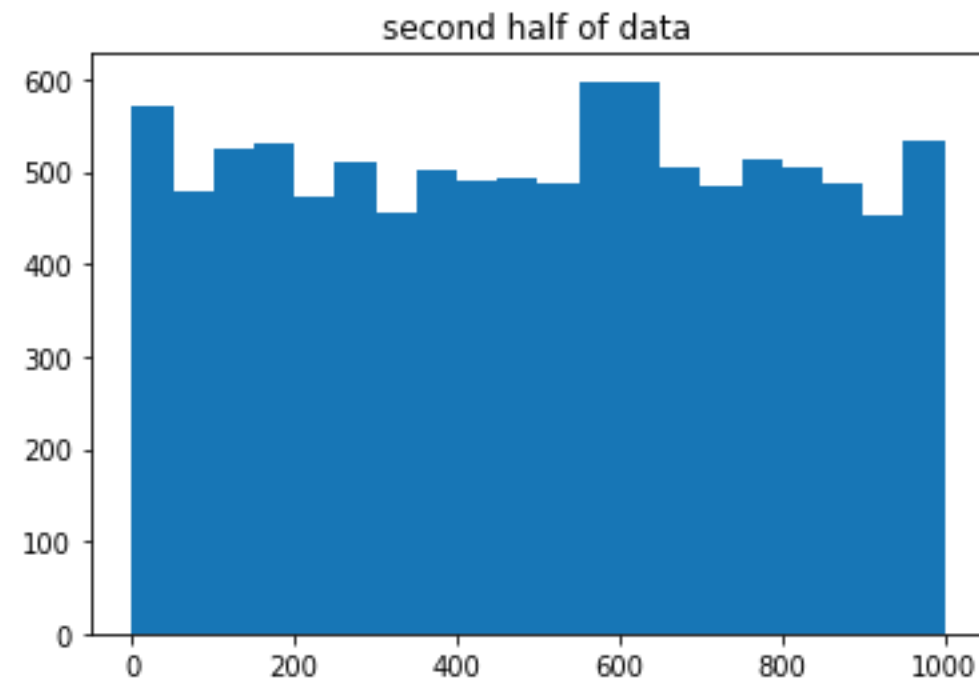
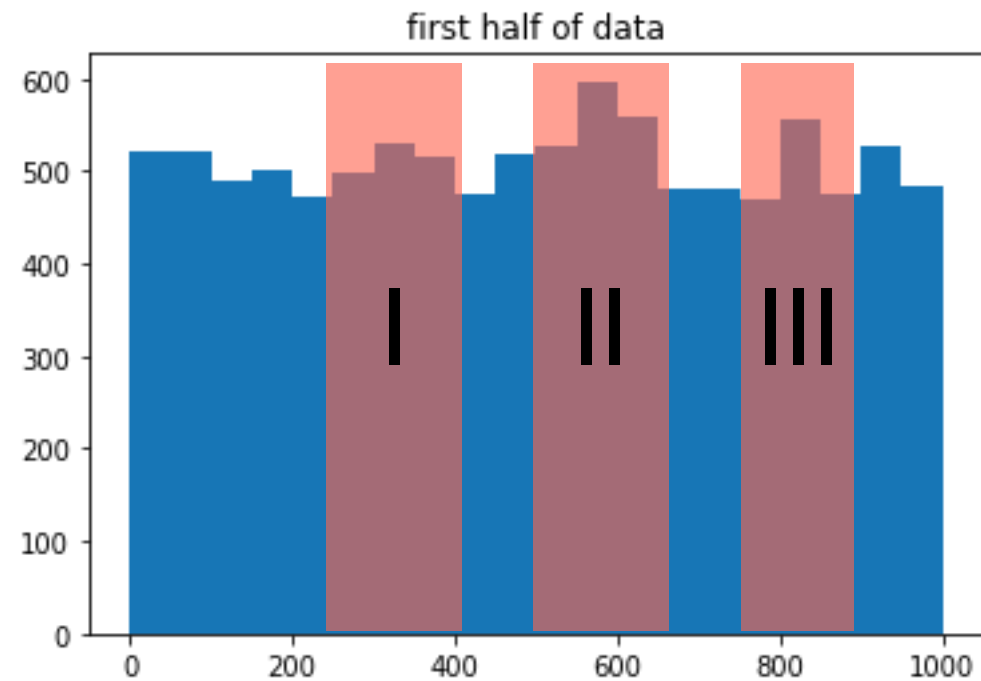
Look elsewhere effect



Simple way to mitigate LEE:

- Divide data in two
- Look for discrepancies in first half

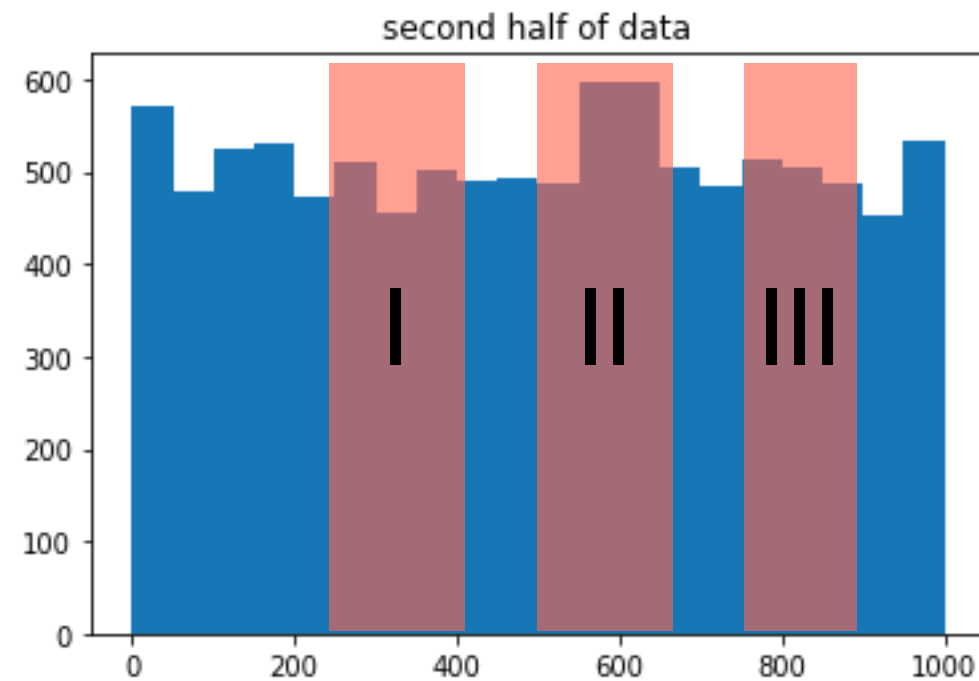
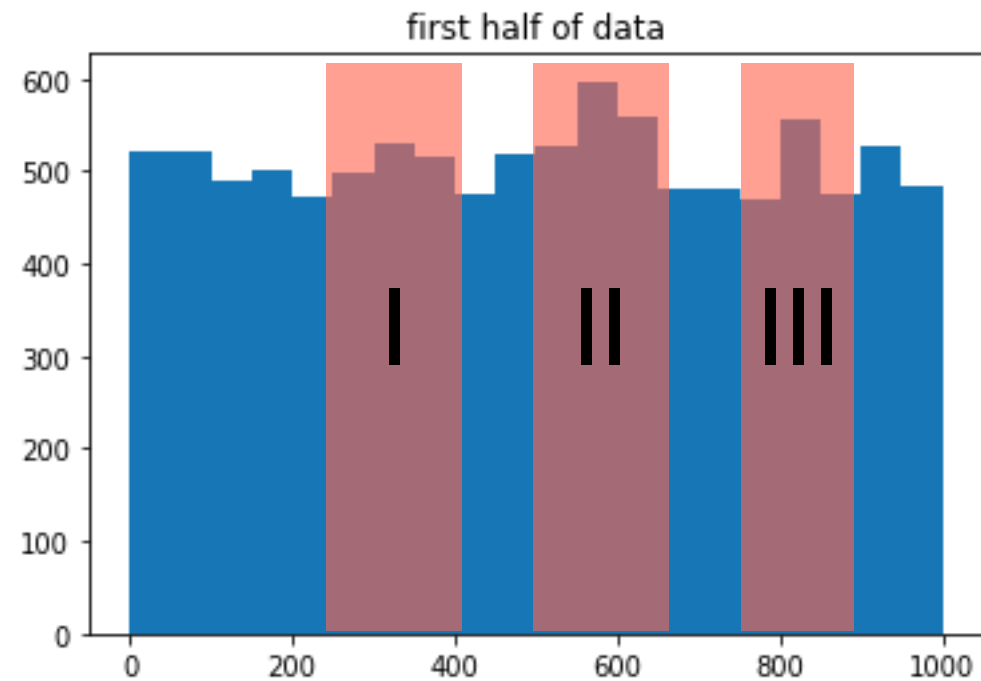
Look elsewhere effect



Simple way to mitigate LEE:

- Divide data in two
- Look for discrepancies in first half
- Fix regions of interest around these discrepancies and restrict search to these in the second half

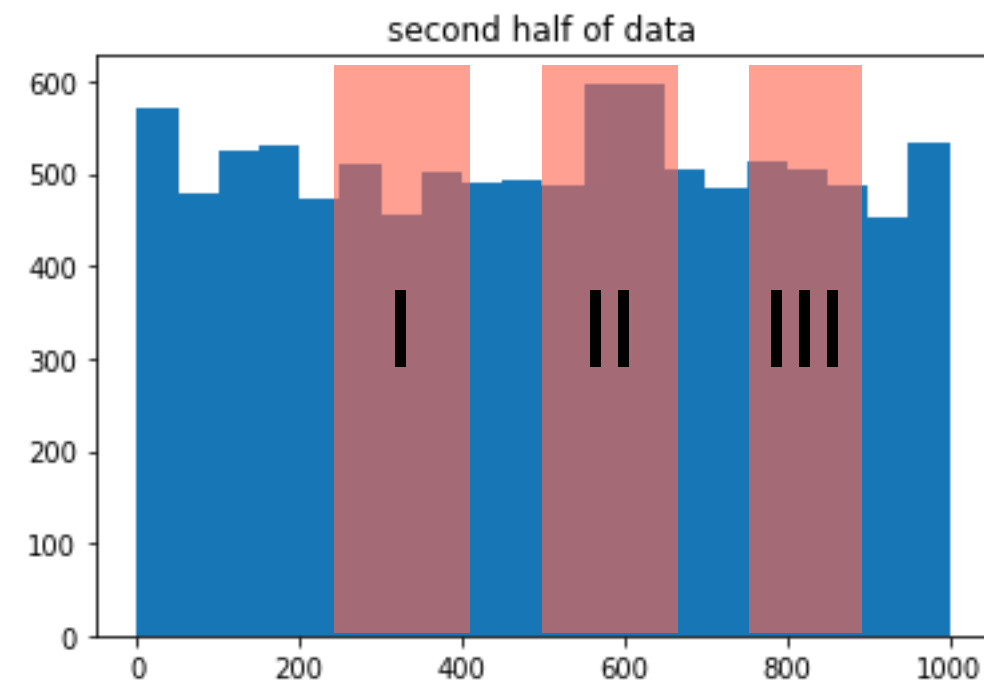
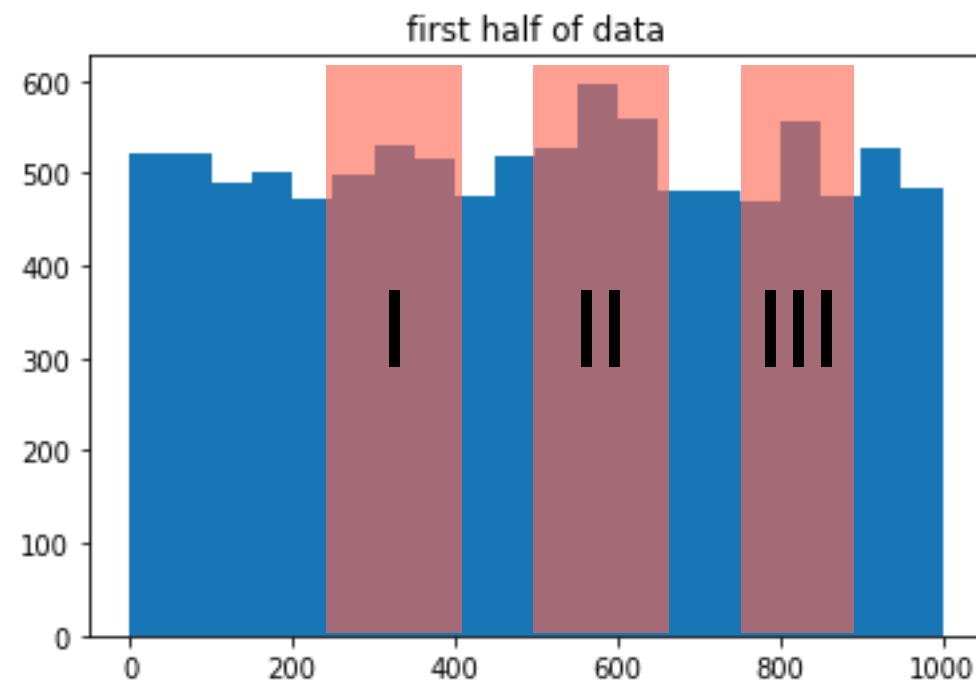
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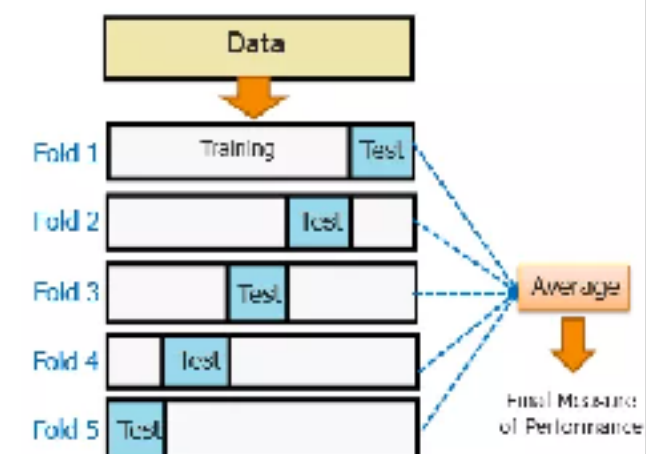
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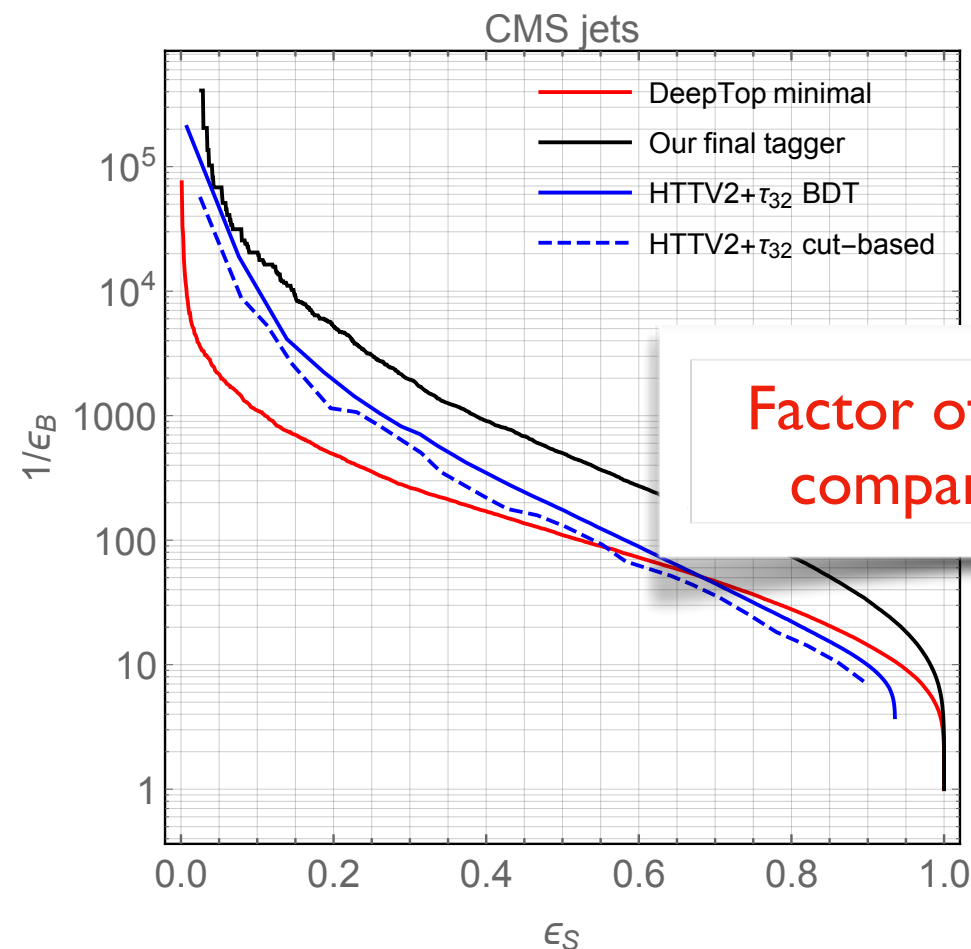
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- An even better approach: k-fold cross validation, see e.g. [1805.02664](#)



Sub-optimal signal/background discrimination

Can do much better than comparing 1D histograms.

Using **deep learning**, can leverage the entire high-dimensional phase space to separate signals from background!



Factor of 3-5 improved performance compared to cut-based methods!

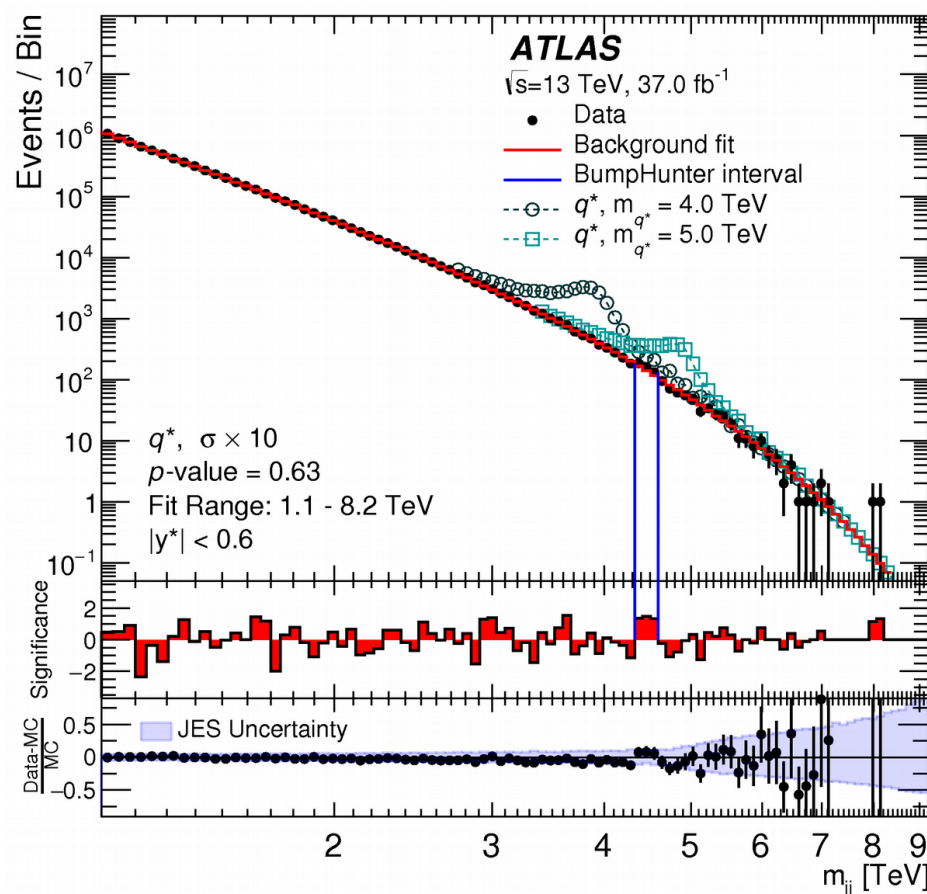
From Macaluso & DS 1803.00107
top tagging with CNNs

See D'Agnolo & Wulzer 1806.02350: train DNN to distinguish data from bg MC.
Optimal version of existing general searches

Over-reliance on simulation for backgrounds

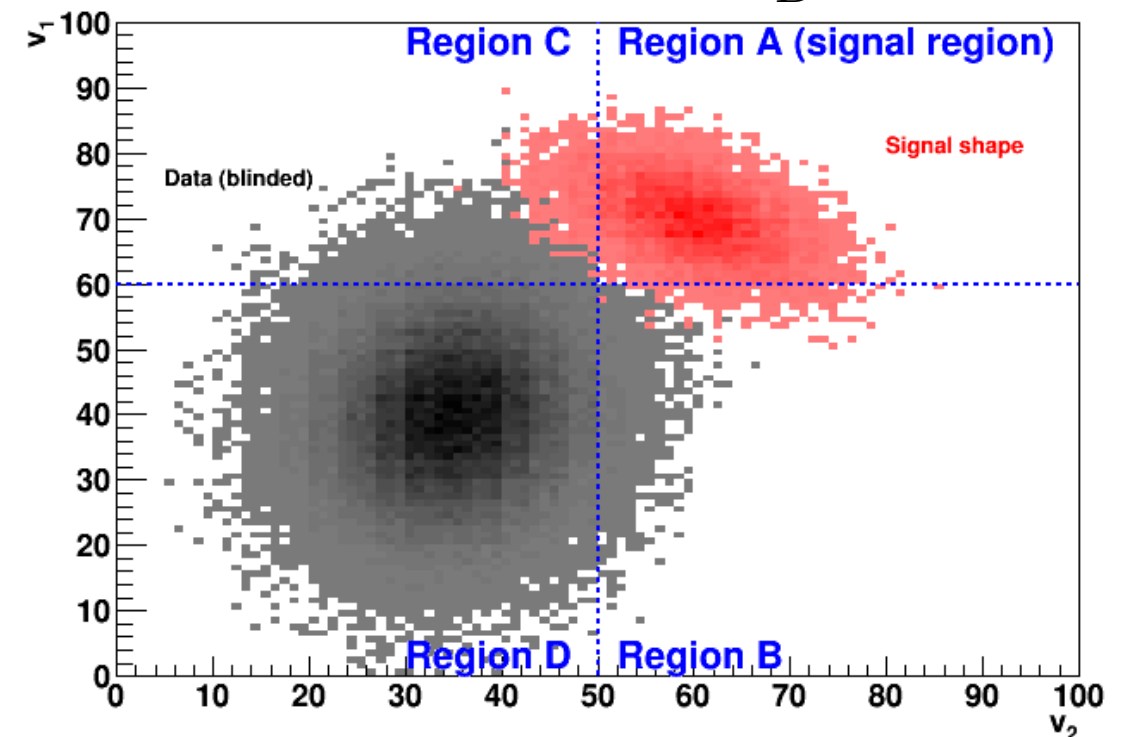
Existing approaches are largely signal model independent, but not background model independent.

To achieve full model independence, we would like to predict backgrounds directly from the data.



Bump hunt with sideband fit

$$N_A = N_B \times \frac{N_C}{N_D}$$



ABCD method

Unsupervised ML for NP searches

Unsupervised ML for NP searches

We want to search for new physics in the data in a signal **and** background independent way.

Cannot use simulation or ground-truth labels. We need **unsupervised ML**.

Two main approaches proposed so far:

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- CWoLa hunting
Collins, Howe & Nachman 1805.02664, 1902.02634

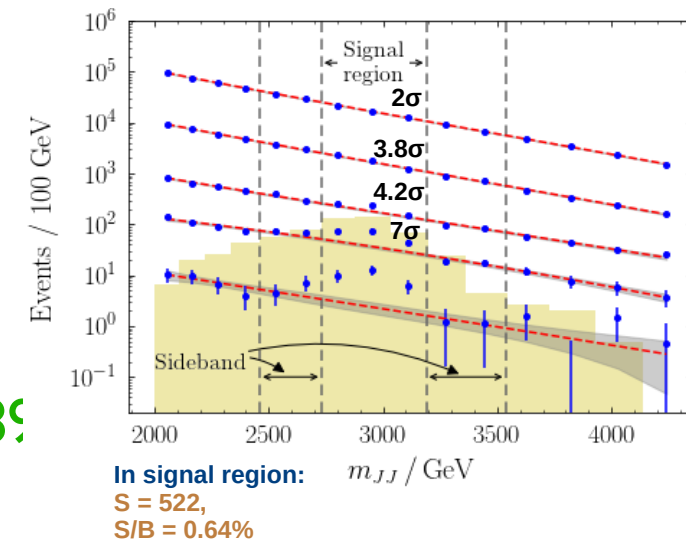
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Farina, Nakai & DS [1808.08992](#); Heimel et al [1808.08992](#)



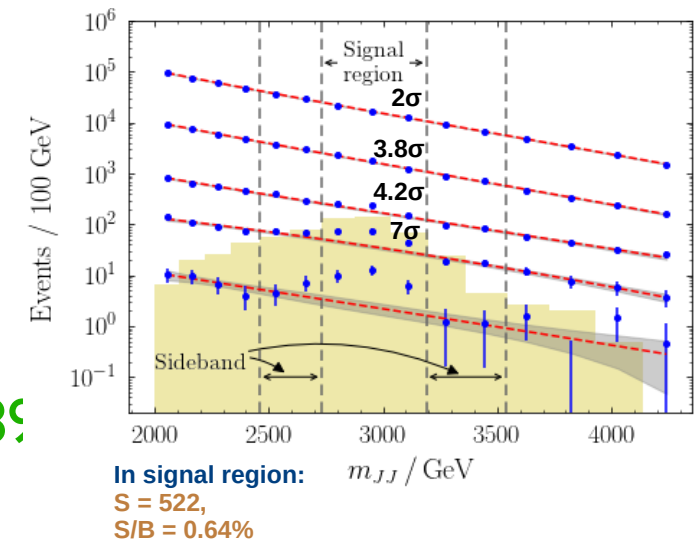
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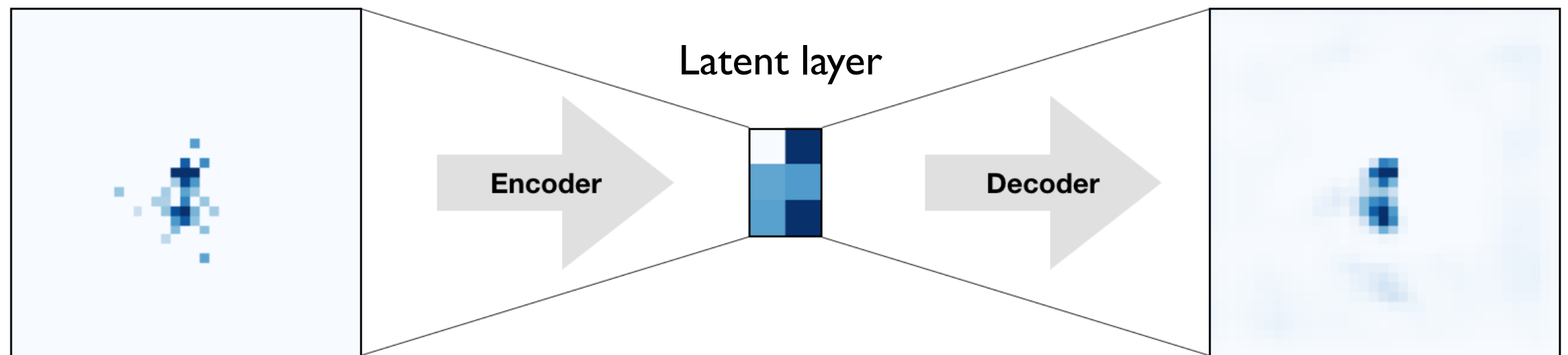
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- Autoencoders
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There are surely more methods waiting to be invented!

Searching for NP with deep autoencoders

Heimel et al I808.08979; Farina, Nakai & DS I808.08992



An autoencoder maps an input into a “latent representation” and then attempts to reconstruct the original input from it.

The encoding is lossy, so the reconstruction is not perfect.

Many real world applications of autoencoders, including anomaly detection, fraud detection, denoising, compression, generation, density estimation

See also:

Hajer et al “Novelty Detection Meets Collider Physics” I807.10261

Cerri et al “Variational Autoencoders for New Physics Mining at the Large Hadron Collider” I811.10276

Searching for NP with deep autoencoders

Heimel et al I808.08979; Farina, Nakai & DS I808.08992

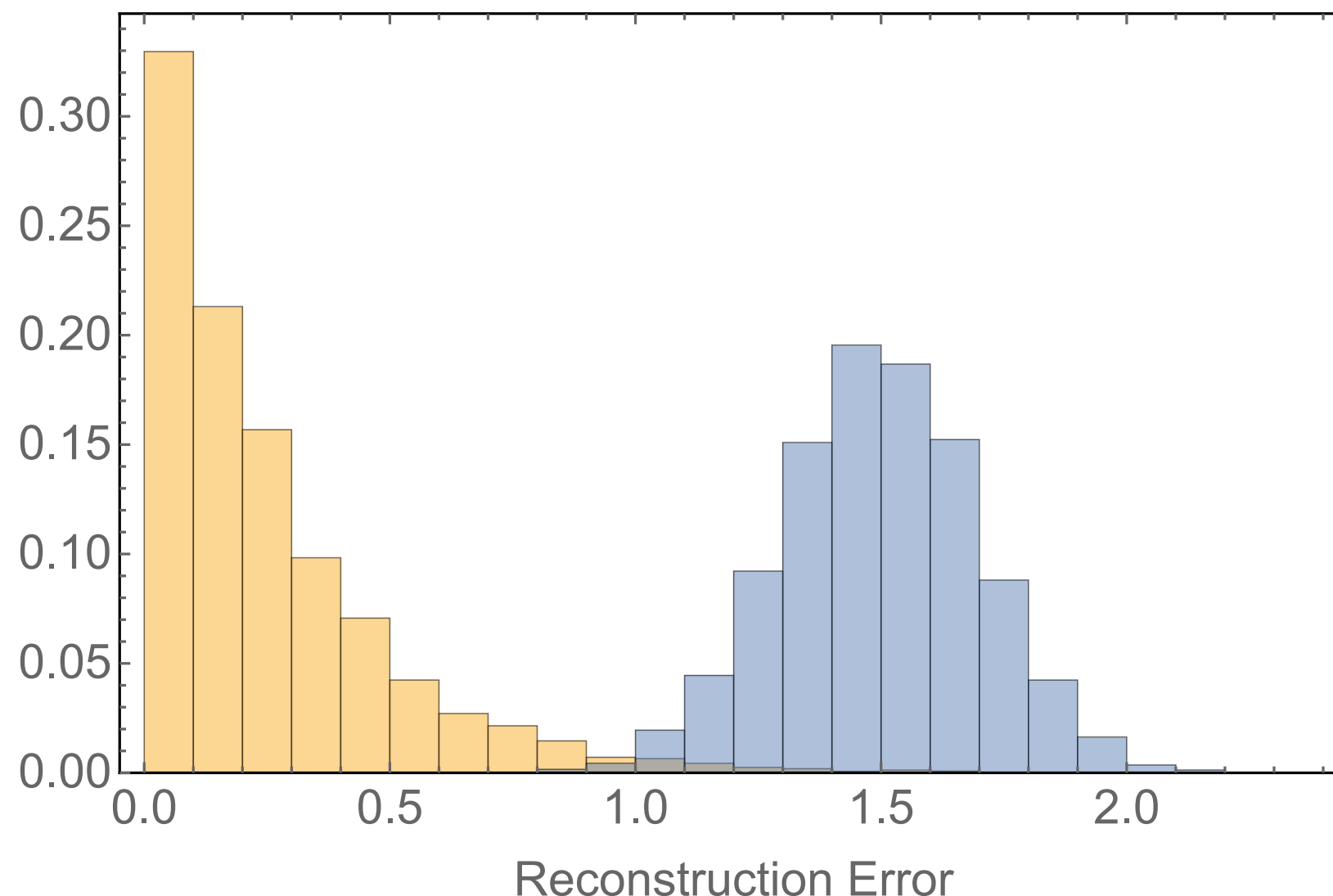
Loss function for autoencoder:
“reconstruction error”

$$L = \frac{1}{N} \sum_{i=1}^N (x_i^{in} - x_i^{out})^2$$

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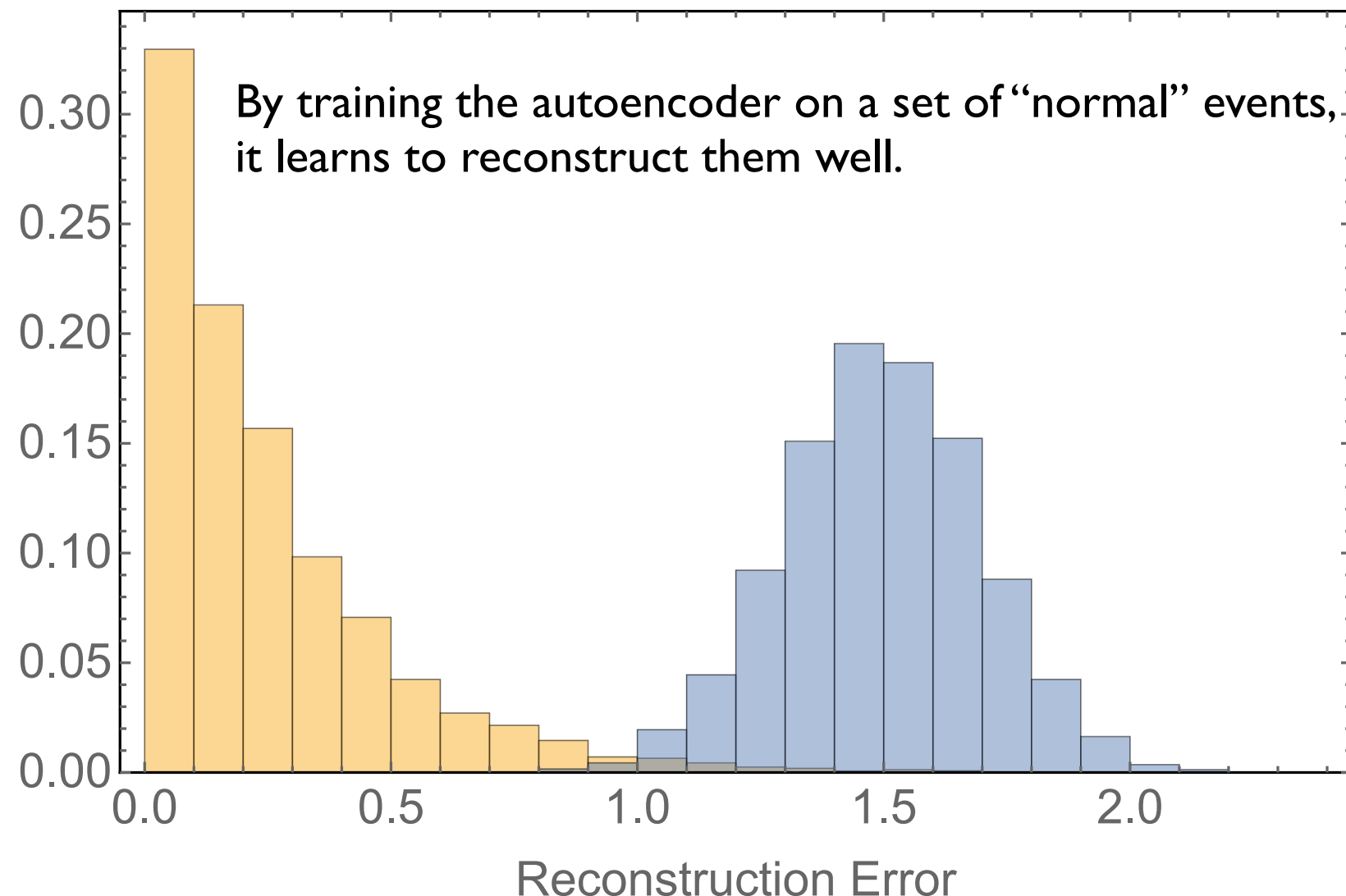
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Searching for NP with deep autoencoders

Heimel et al 1808.08979; Farina, Nakai & DS 1808.08992

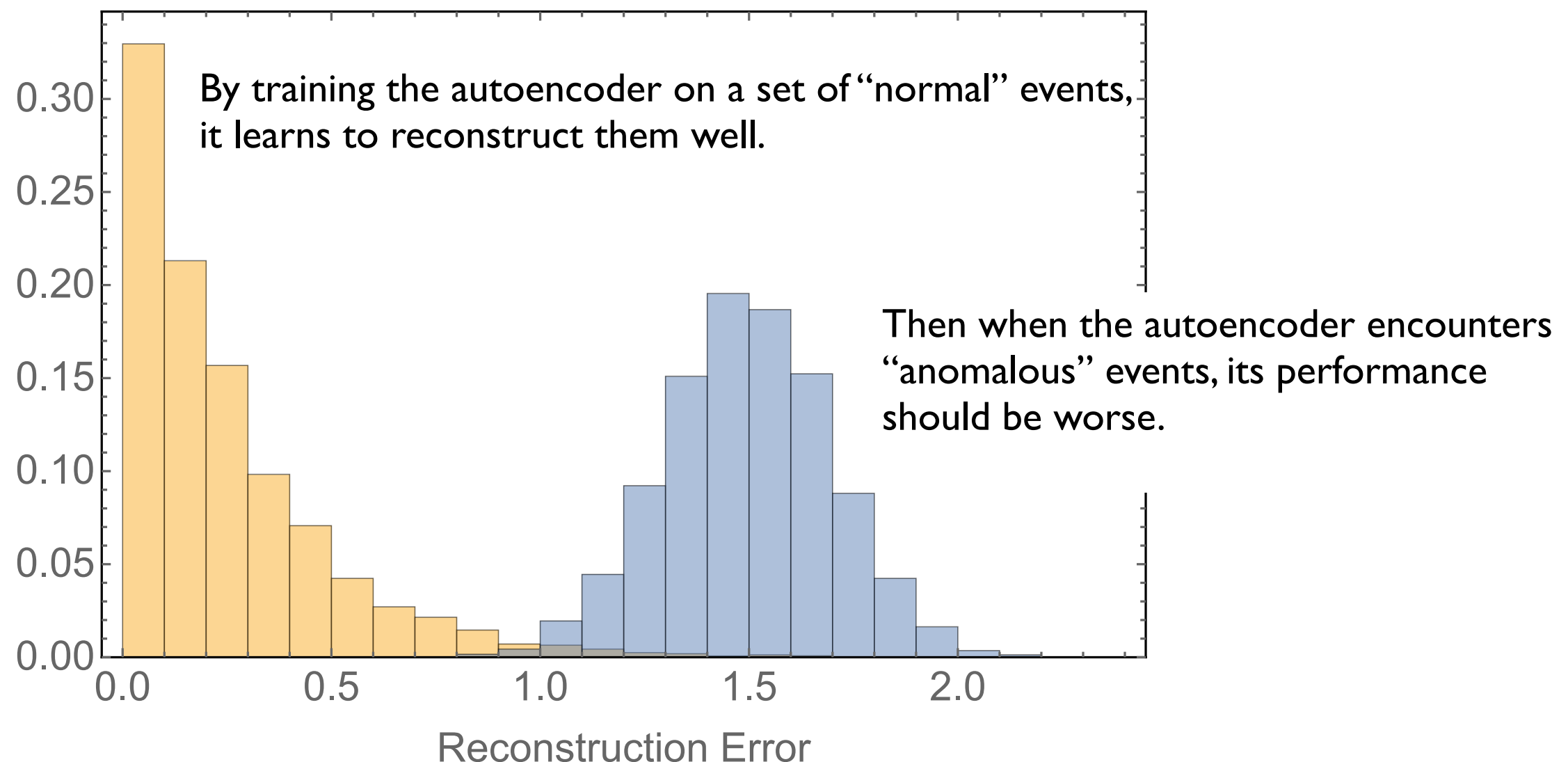
Loss function for autoencoder: $L = \frac{1}{N} \sum_{i=1}^N (x_i^{in} - x_i^{out})^2$
“reconstruction error”



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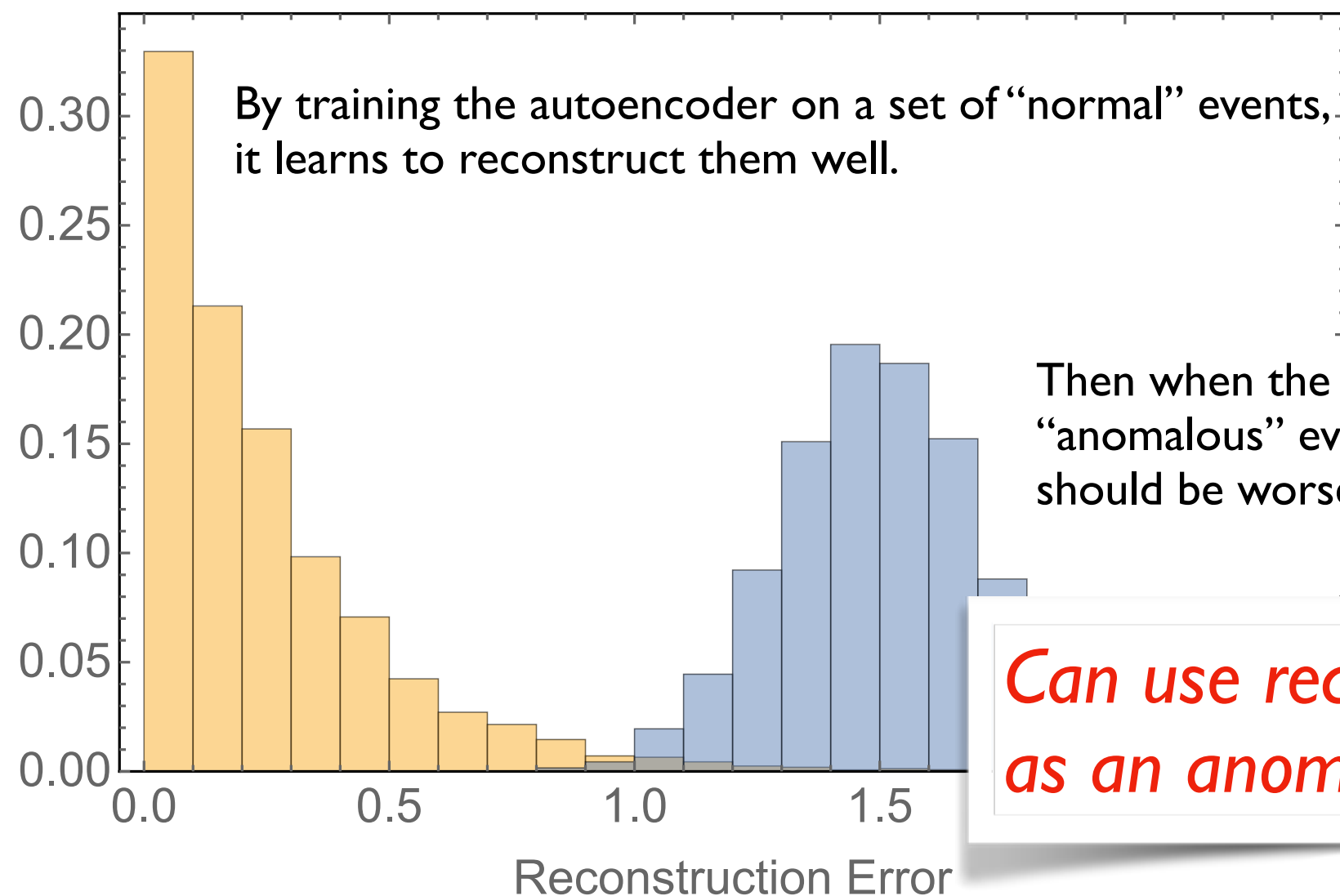
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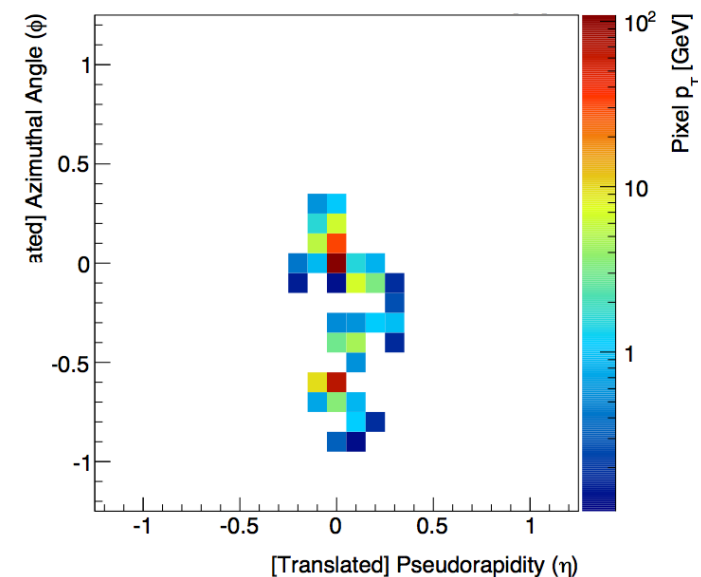
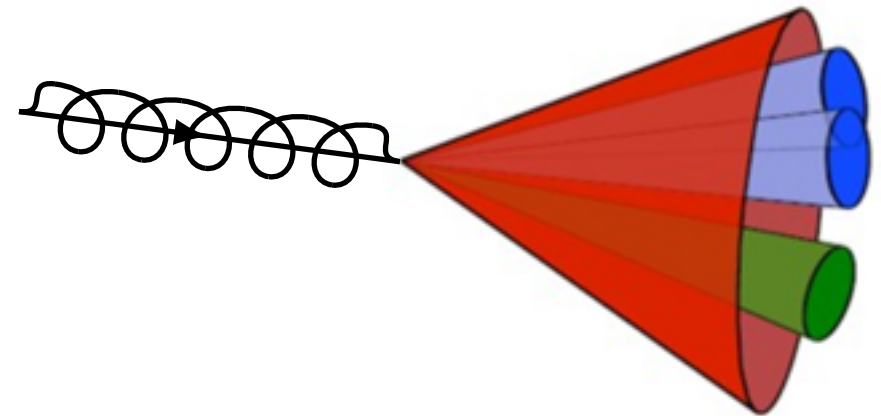
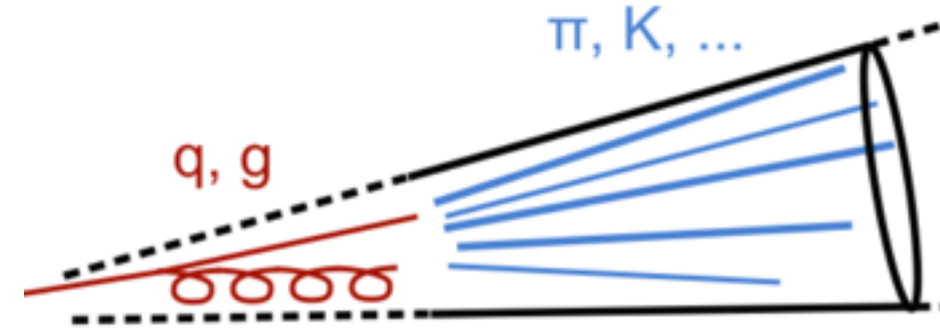
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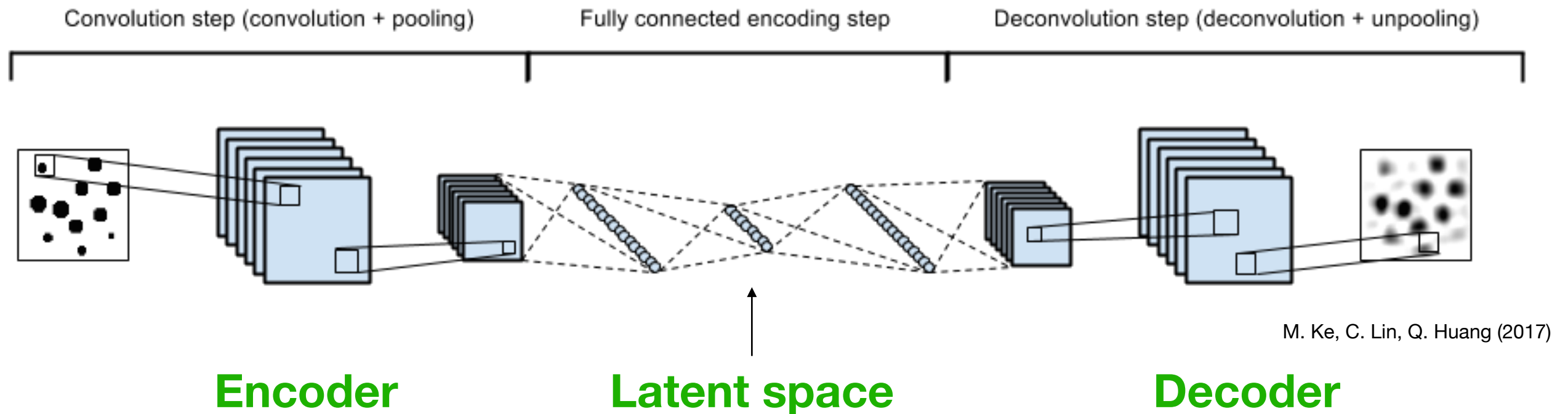
Can use reconstruction error as an anomaly threshold!

Sample definitions

- Background: QCD jets
(p_T : 800-900 GeV, $|\eta| < 1$, anti-kt $R=1$)
- Signals:
 - All-hadronic tops
 - 400 GeV gluinos decaying via RPV
- We formed jet images in η and ϕ with a pixel resolution and intensity given by the calorimeter towers.



Autoencoder architecture



I 28C3-MP2-I 28C3-MP2-I 28C3-32N-6N-32N-I 2800N-I 28C3-US2-I 28C3-US2-I C3

Our primary autoencoder used **convolutional neural networks (CNNs)** for encoding and decoding the jet images.

We also considered autoencoders based on PCA and simple DNNs.

Many more architectures are possible.

Choosing the latent dimension

Should choose the latent dimension in an unsupervised manner
(ie without optimizing on a specific signal)

- d too large $\rightarrow$ autoencoder becomes identity transform
- d too small $\rightarrow$ autoencoder cannot learn all the features

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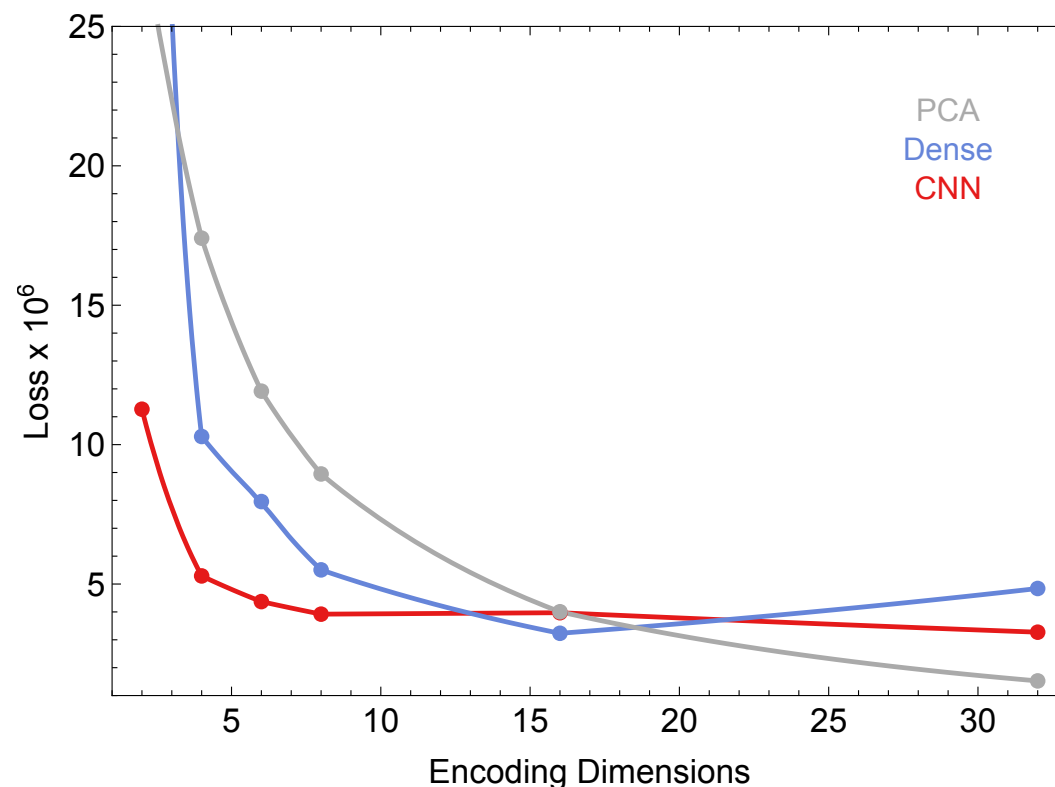
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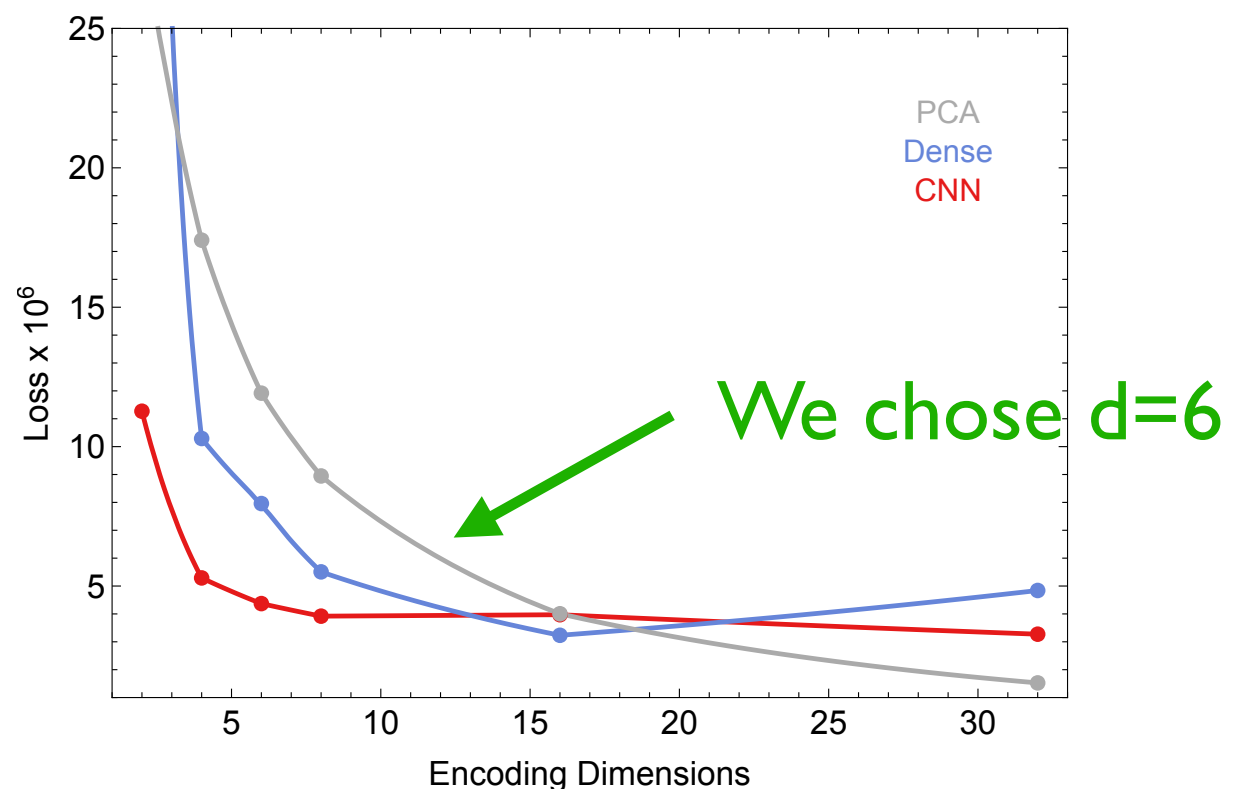


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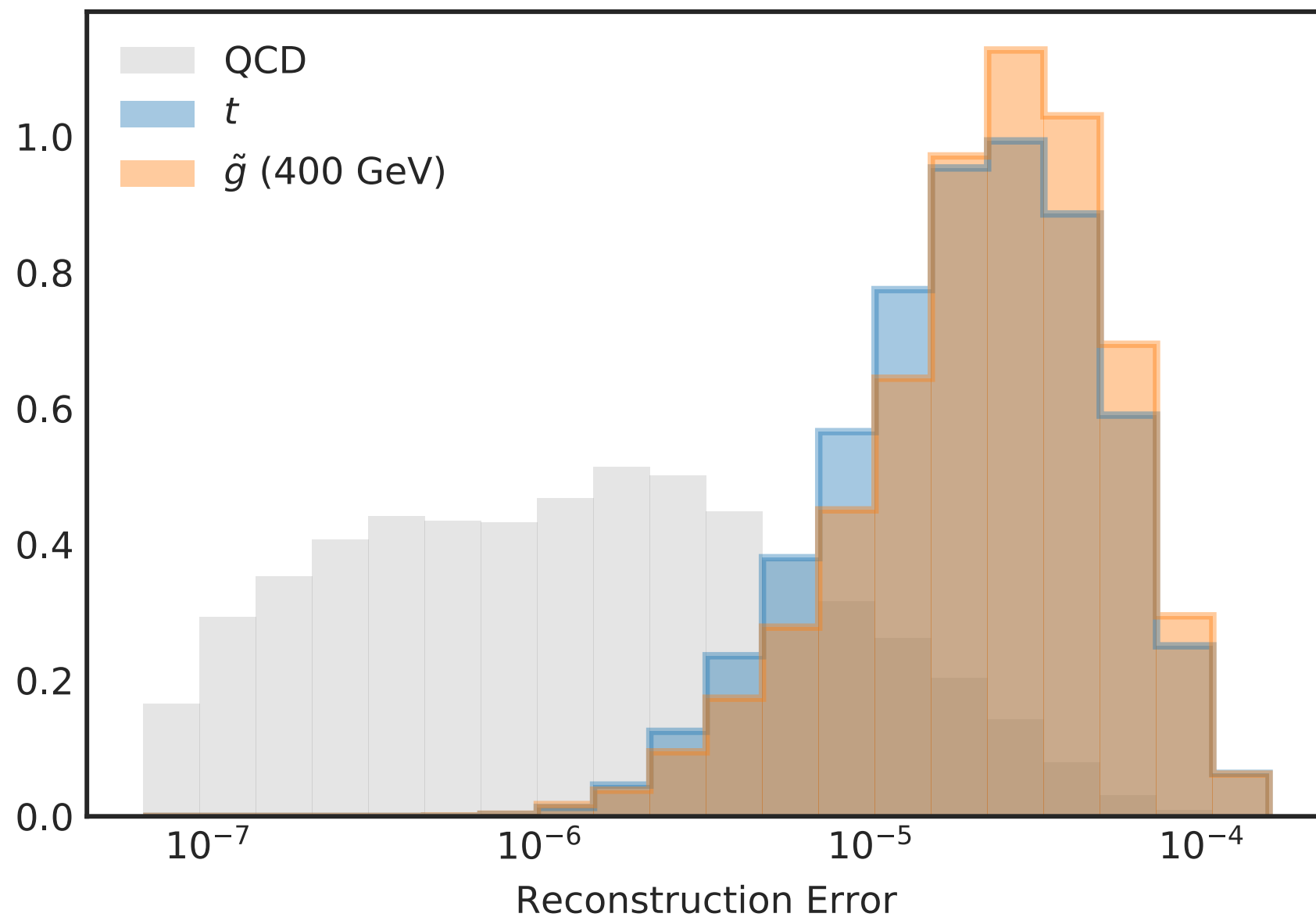
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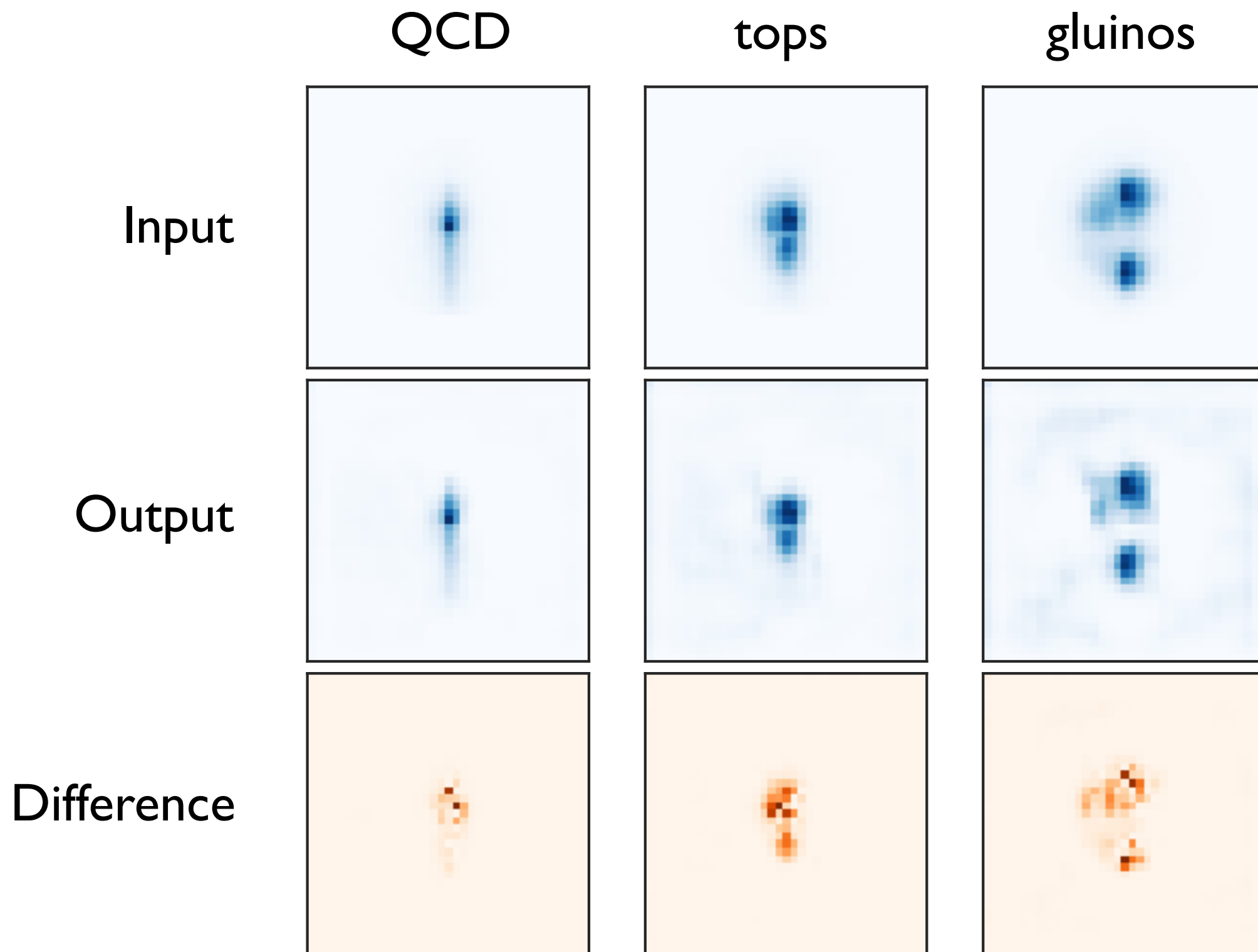
Performance: weakly supervised mode

Train the AE on QCD backgrounds only.



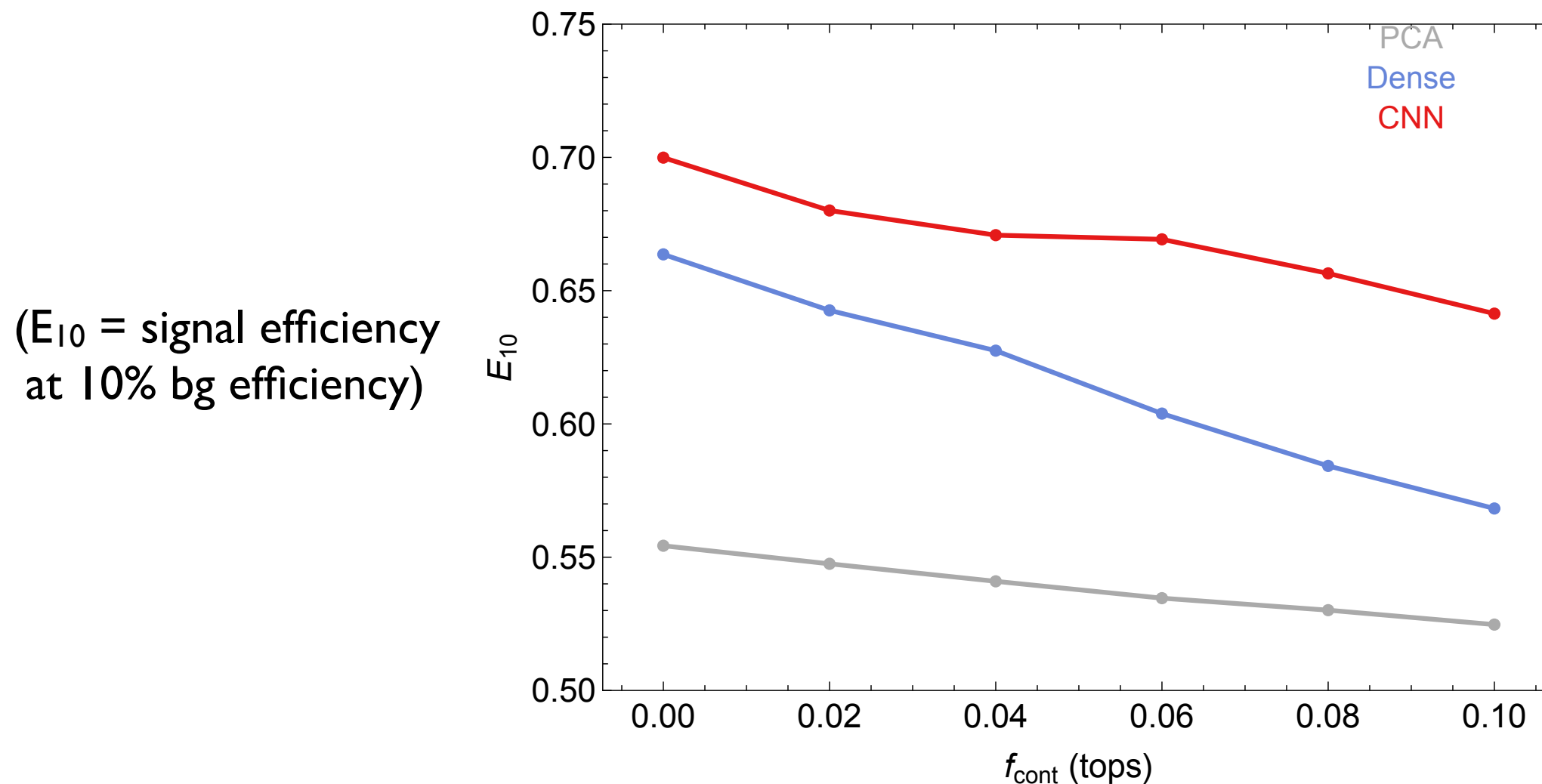
It works as an anomaly detector!

Performance: weakly supervised mode



Fully unsupervised mode

Can also train on QCD background “contaminated” with a small fraction of signal. **This could be representative of actual data.**



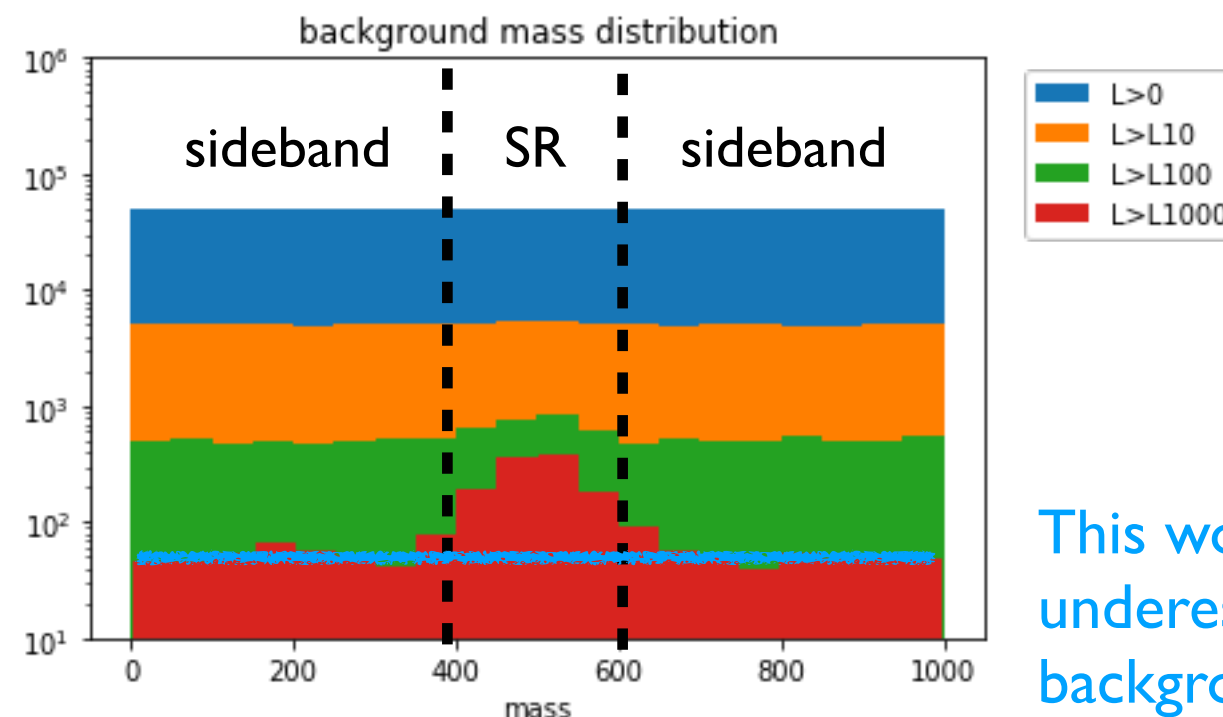
Performance of AE robust even up to 10% contamination!

Background estimation

Finally, background estimation. Remember: unlike past attempts at model-independent searches, we want it to be data-driven!

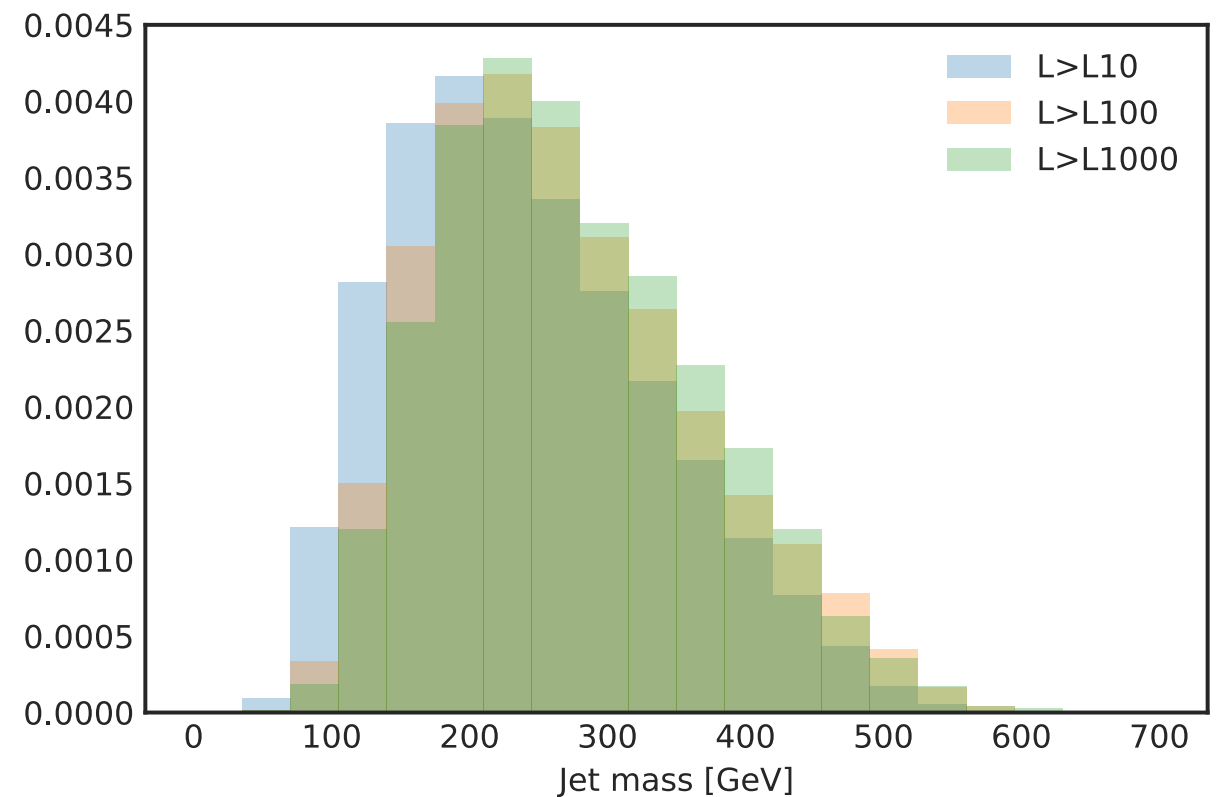
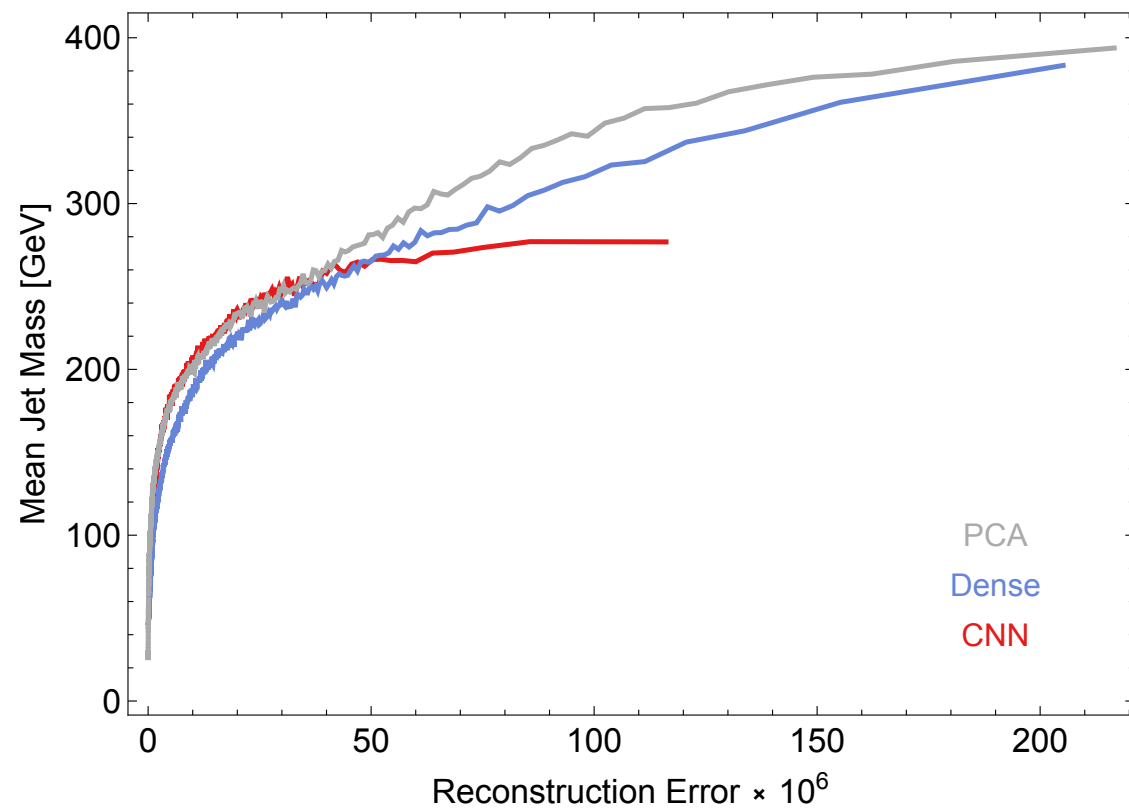
One idea: **combine autoencoder with a bump hunt in jet mass.**
Estimate backgrounds using sidebands in mass.

Only works if cutting on reconstruction error does not sculpt the mass distribution of the background!



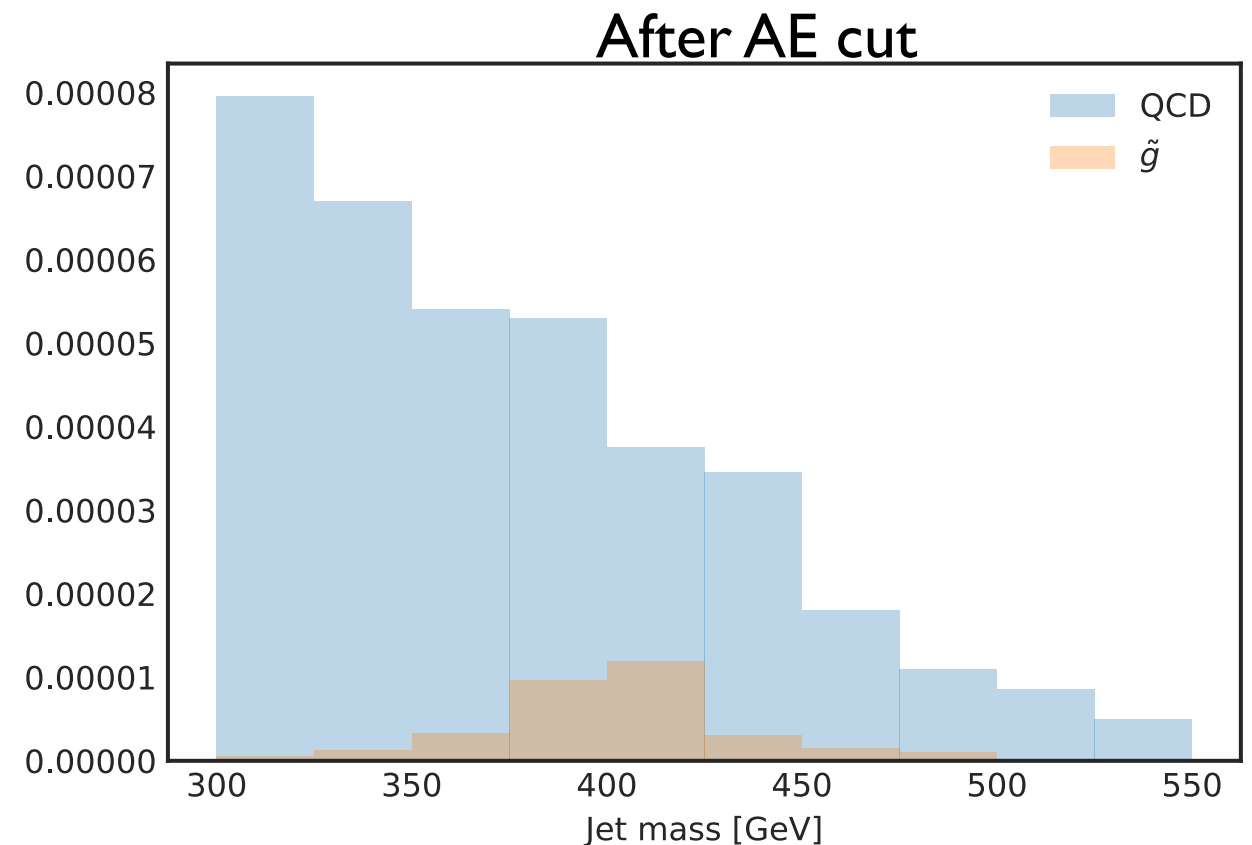
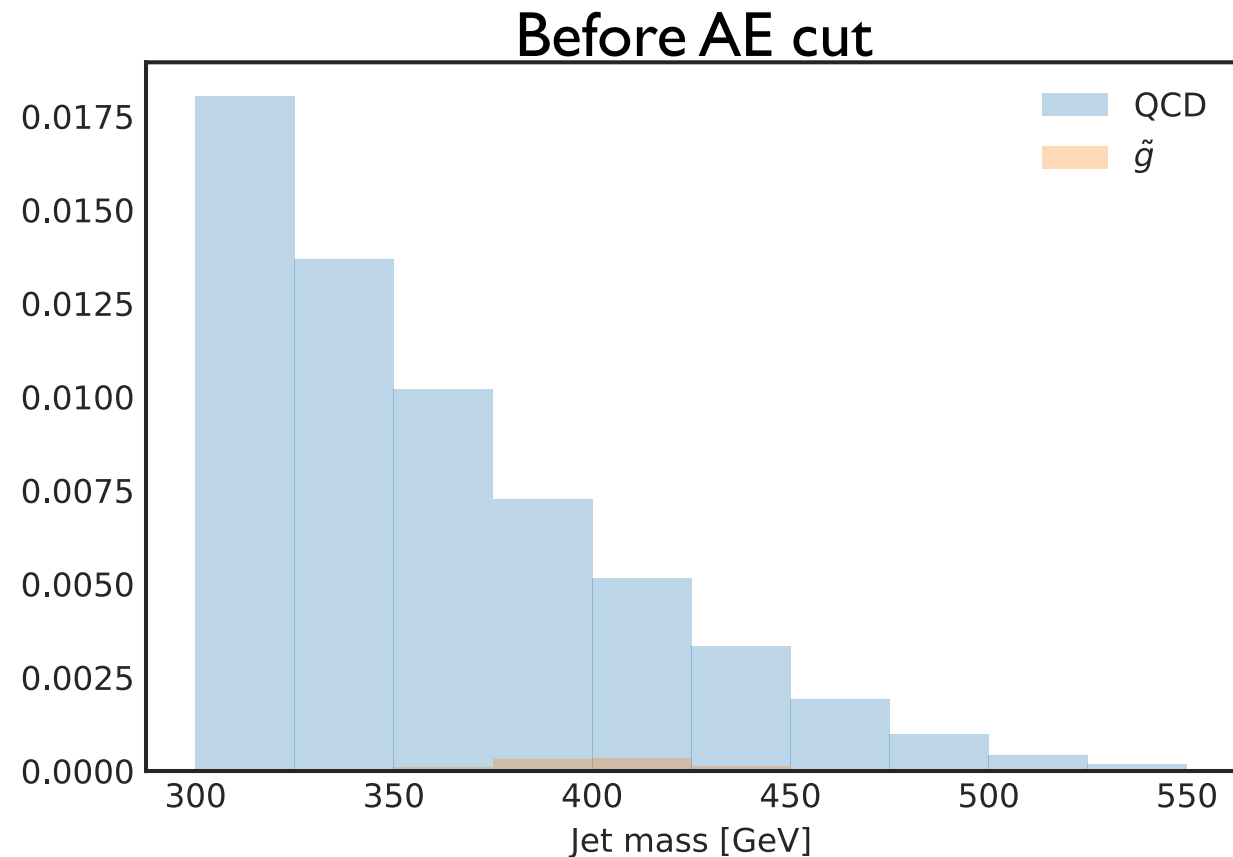
This would greatly underestimate the background in the SR

Bump hunt with autoencoder



We find empirically that the background jet mass distribution is fairly stable against cuts on CNN AE reconstruction loss above ~250 GeV.

Bump hunt with autoencoder



Train directly on data that contains 400 GeV gluinos.

Use the AE to clean away QCD jets.

Enhance the significance of the bump hunt! (improve S/B by factor of ~ 6)

Could really discover new physics this way!

Autoencoder with explicit decorrelation

A more controlled approach to mass decorrelation would be to explicitly penalize correlations in the training of the autoencoder.

One promising method: autoencoder with adversarial decorrelation (Heimel et al 1808.08979; based on 1611.01046, 1703.03507 and the idea behind GANs)

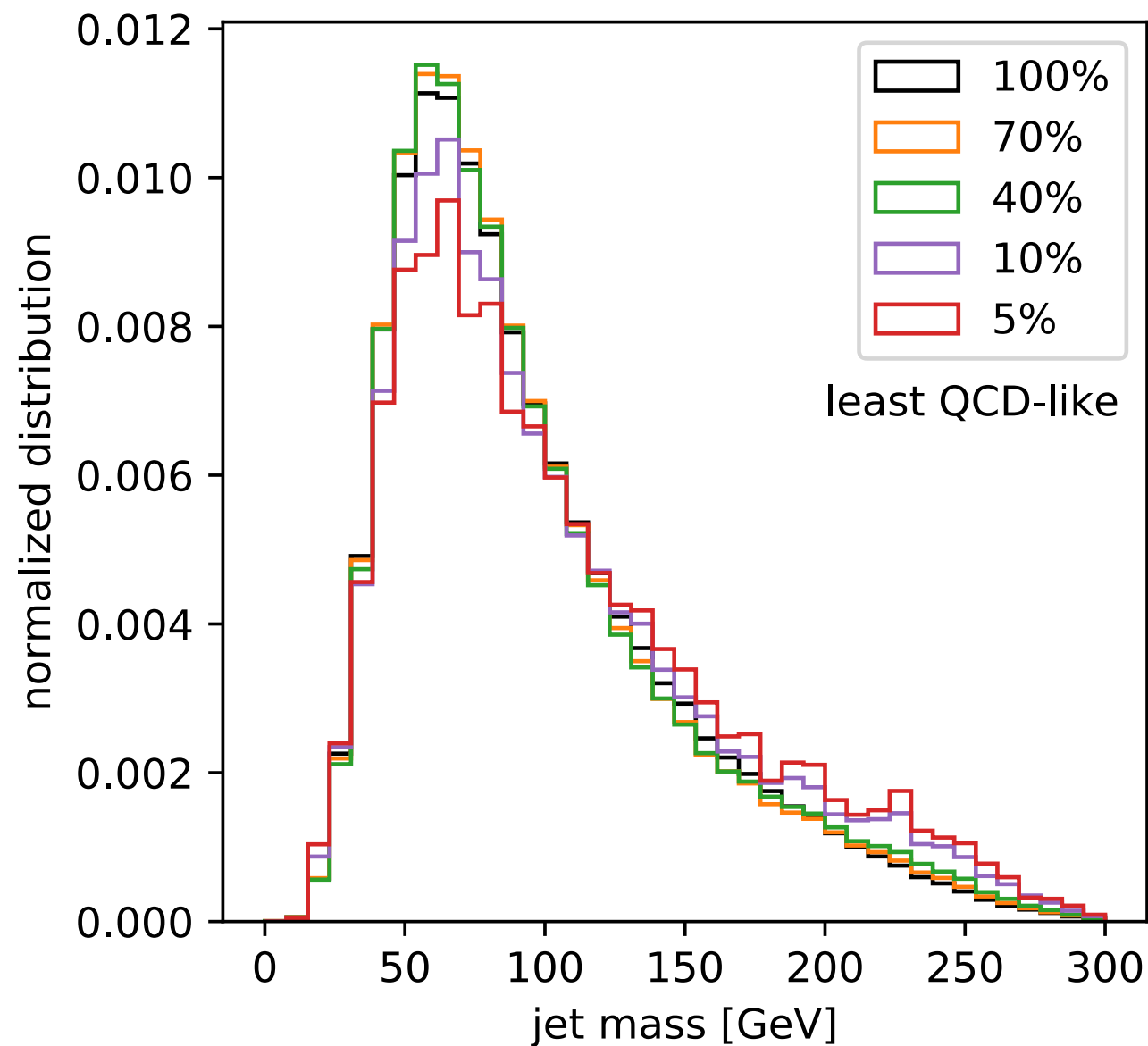
- Introduce a second NN, **the adversary**, that tries to predict the mass from the reconstruction loss.
- Penalize the total loss function when the adversary does well.

$$L_{adv} = \sum_i (f_{adv}(L_{AE}(x_i)) - m_i)^2$$

$$L_{tot} = L_{AE} - \lambda L_{adv}$$

Autoencoder with explicit decorrelation

Heimel et al 1808.08979



Current implementation needs QCD backgrounds for decorrelation.
Can generalize to fully unsupervised case?

Alternatives to adversaries

Adversaries are notoriously **tricky to train** — saddle point optimization

$$\min_{\theta_{\text{clf}}} \max_{\theta_{\text{adv}}} L_{\text{clf}}(y(\theta_{\text{clf}})) - \lambda L_{\text{adv}}(y(\theta_{\text{clf}}), m; \theta_{\text{adv}})$$

Would be great if we could achieve the same performance but with a convex regularizer term

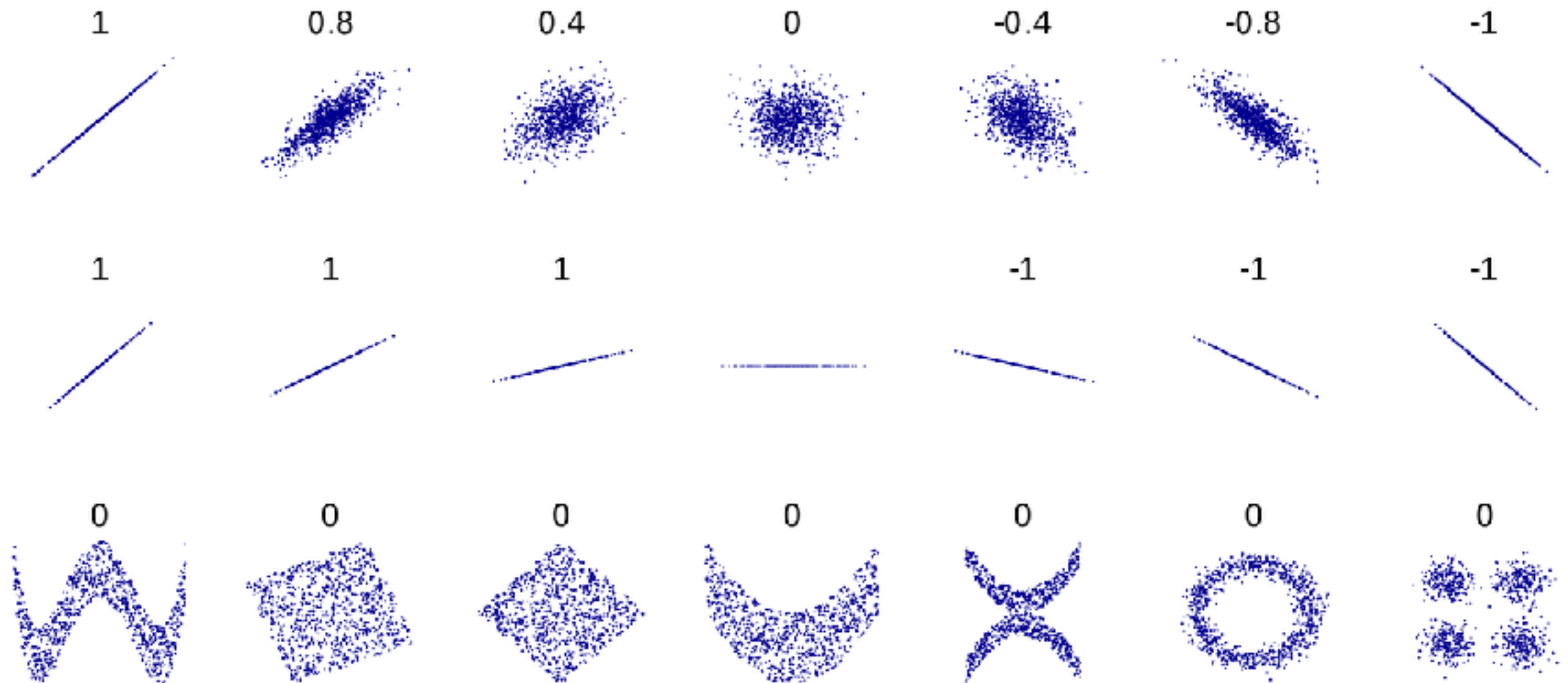
$$\min_{\theta_{\text{clf}}} L_{\text{clf}}(y(\theta_{\text{clf}})) + \lambda C_{\text{reg}}(y(\theta_{\text{clf}}), m)$$

First idea: can we just use Pearson correlation coefficient?

$$C_{\text{reg}} = R(y, m) \propto \sum_i y_i m_i$$

Problem: this only measures linear correlations

Pearson correlation



y and m can be highly correlated yet $R=0$

Distance correlation (“DisCo”)

Work in progress with Gregor Kasieczka

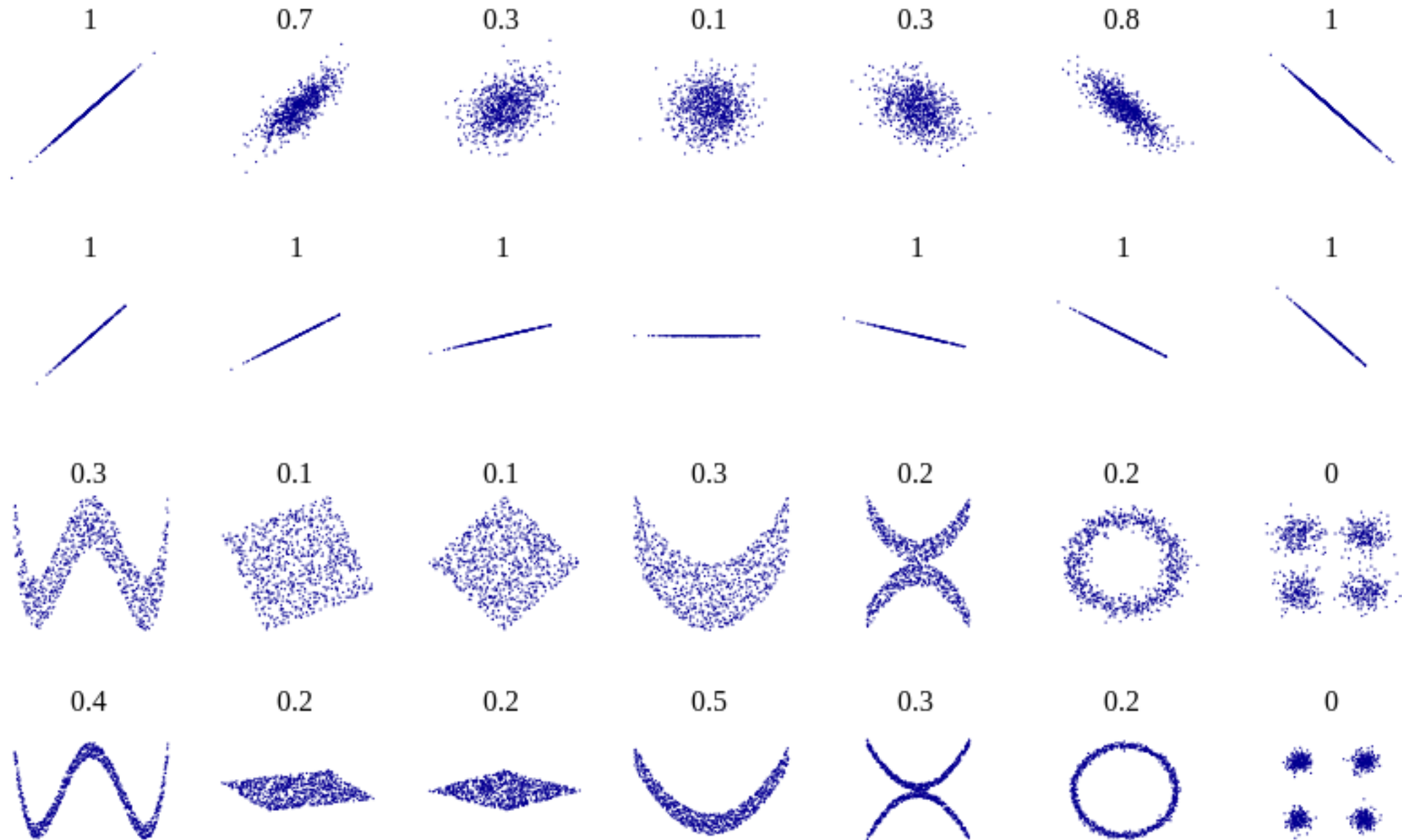
Promising idea: “distance correlation”

(Szekely, Rizzo, Bakirov 2007; Szekely & Rizzo 2009)

$$\text{dCov}^2(X, Y) = \langle |X - X'| |Y - Y'| \rangle + \langle |X - X'| \rangle \langle |Y - Y'| \rangle - 2 \langle |X - X'| |Y - Y''| \rangle$$

- Zero iff X, Y are independent; positive otherwise
- Computationally tractable
- Straightforward sample definition — doesn’t require binning

Distance correlation



Disco is sensitive to nonlinear correlations!

State of the art: ATLAS study of various decorrelation methods in context of boosted W-tagging.

ATLAS Simulation Preliminary

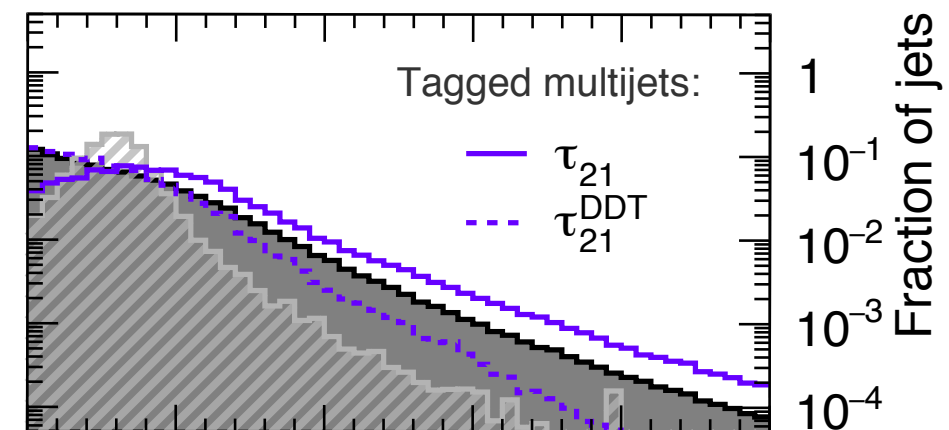
$\sqrt{s} = 13$ TeV, W jet tagging

Cuts at $\varepsilon_{\text{sig}}^{\text{rel}} = 50\%$

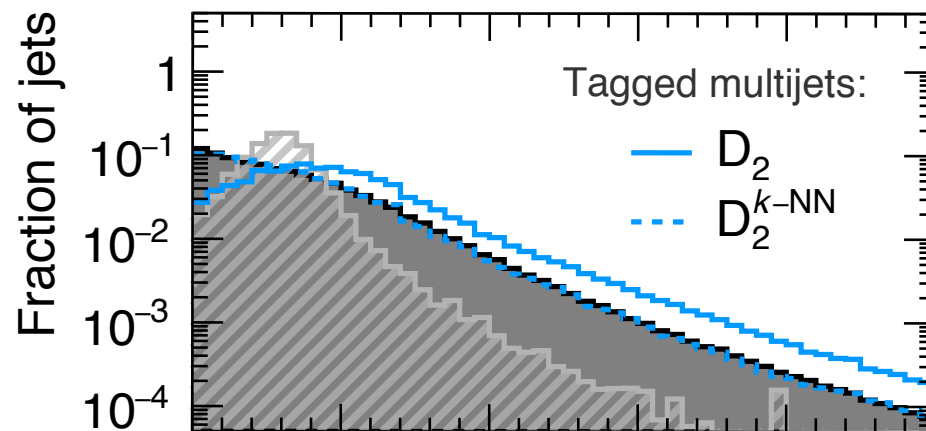
Inclusive selection:

■ Multijets ▨ W jets

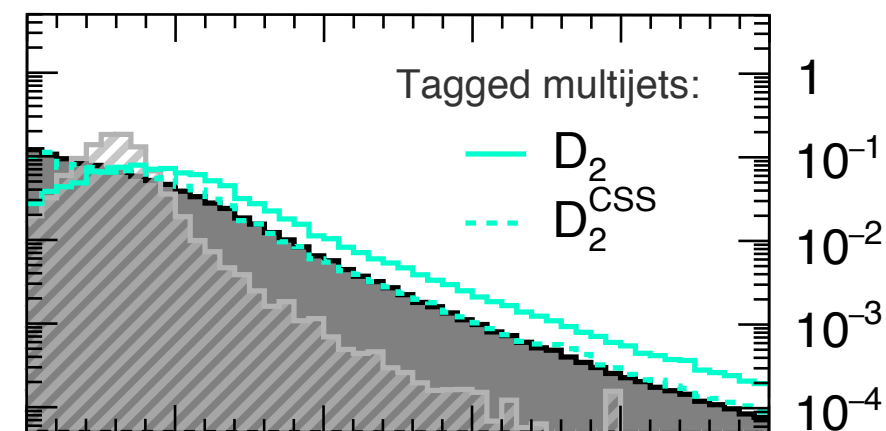
ATL-PHYS-PUB-2018-014



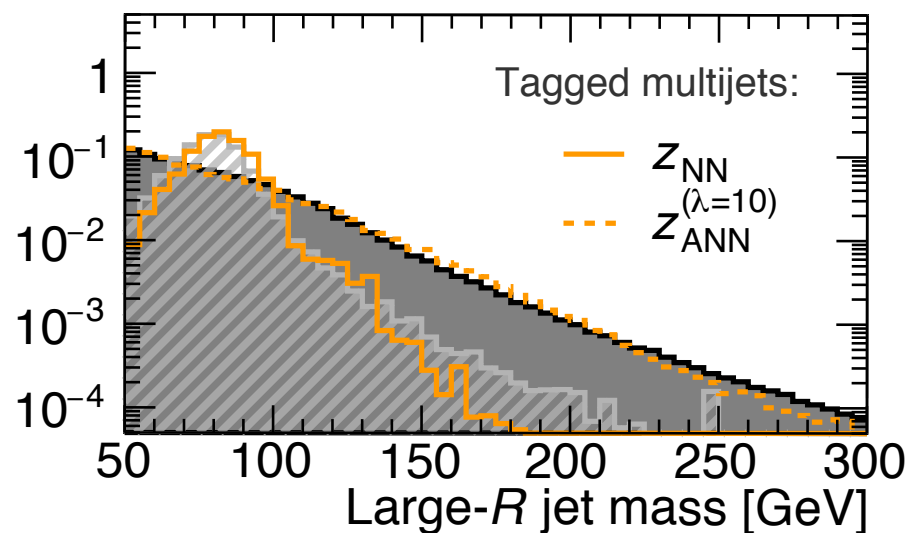
DDT



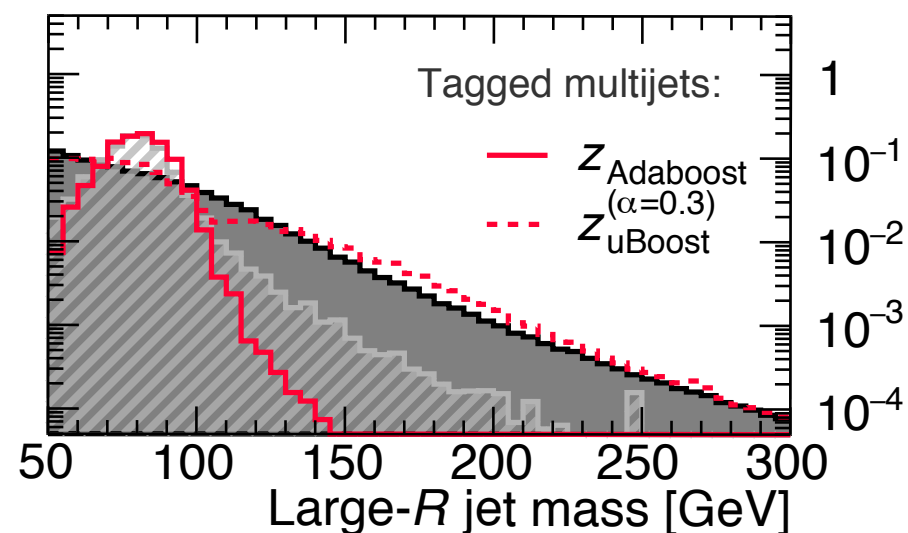
$k\text{-NN}$



CSS



ANN

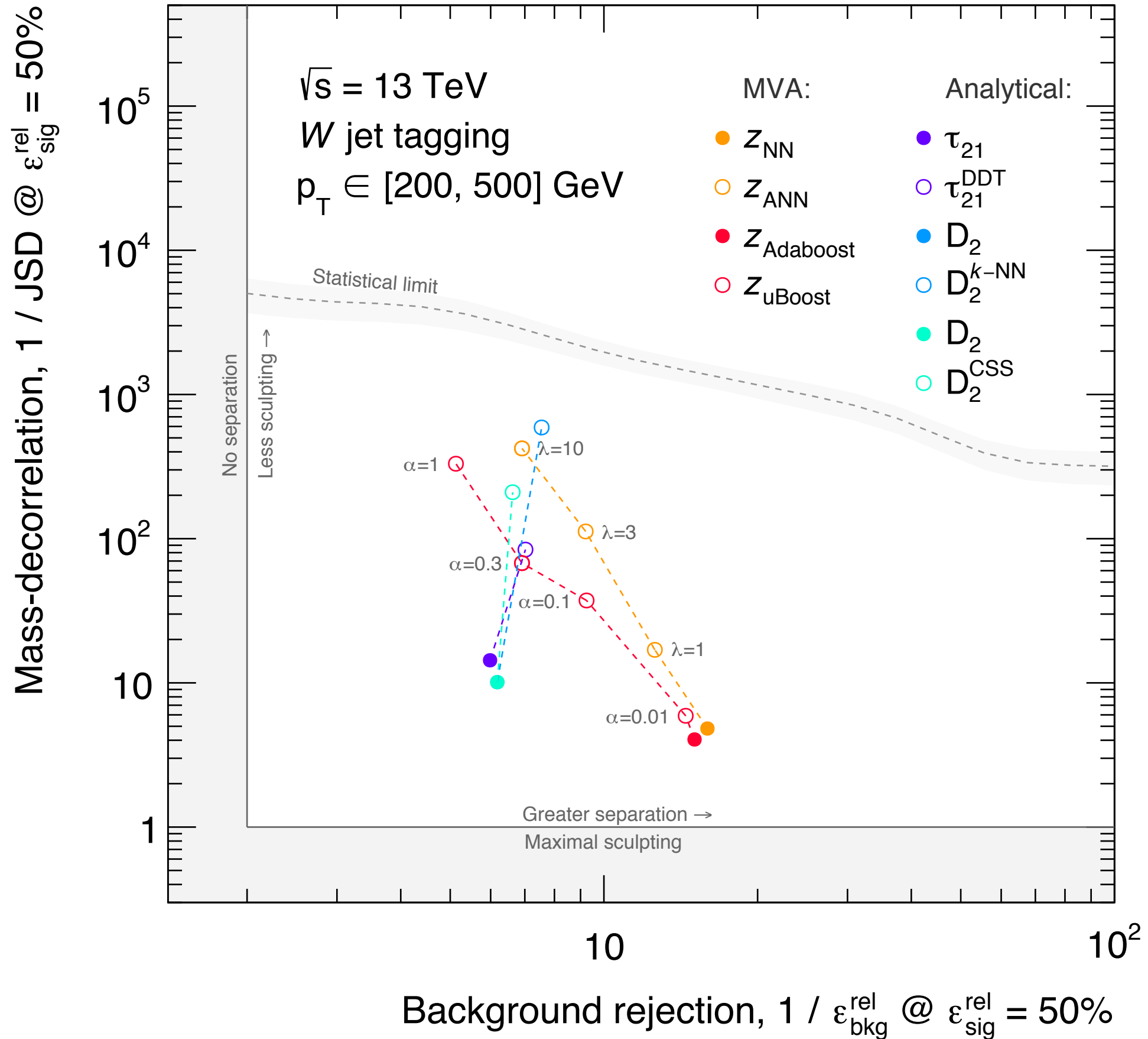


uBoost

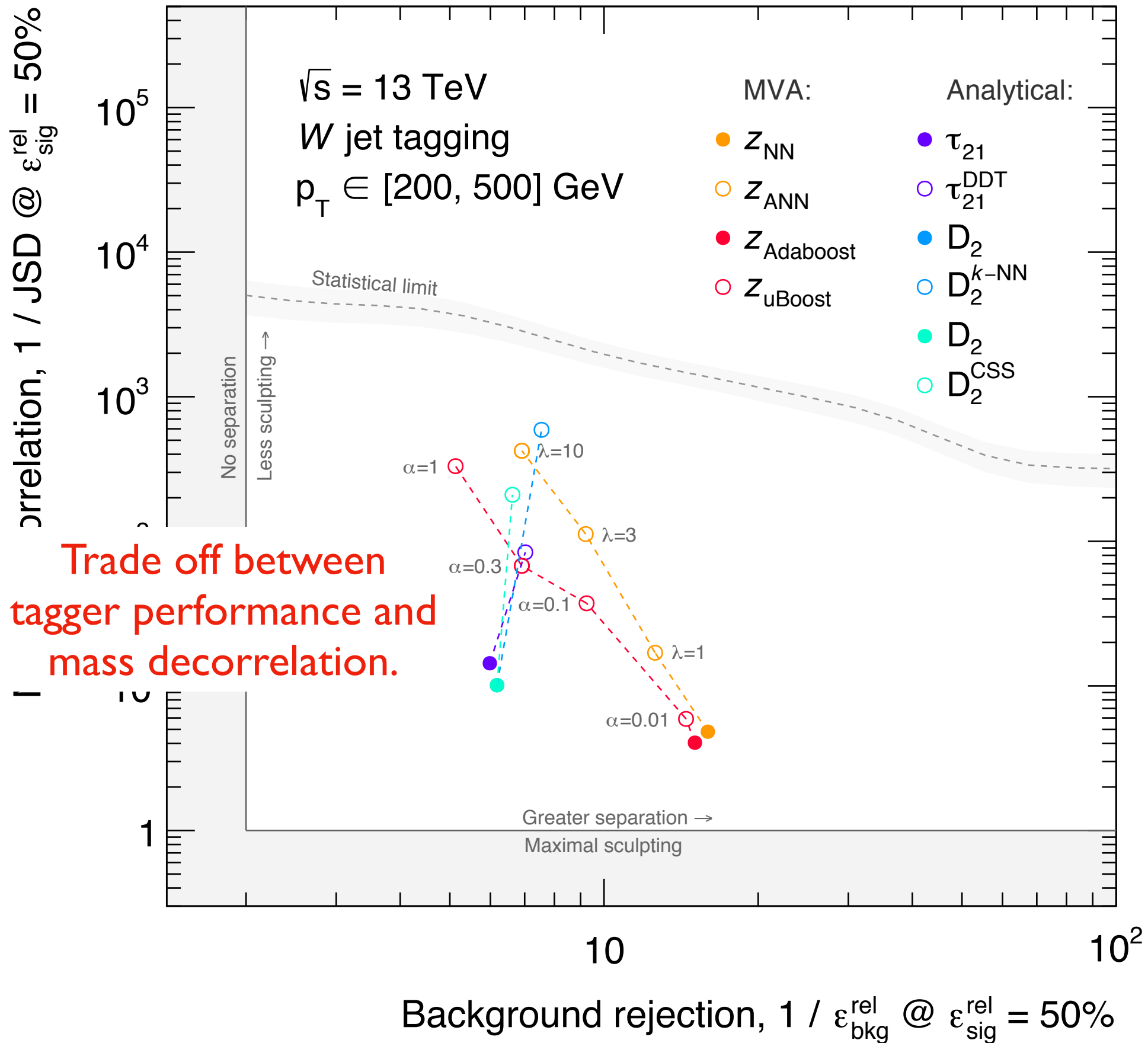
Analytical / single-variable

Multivariate (MVA)

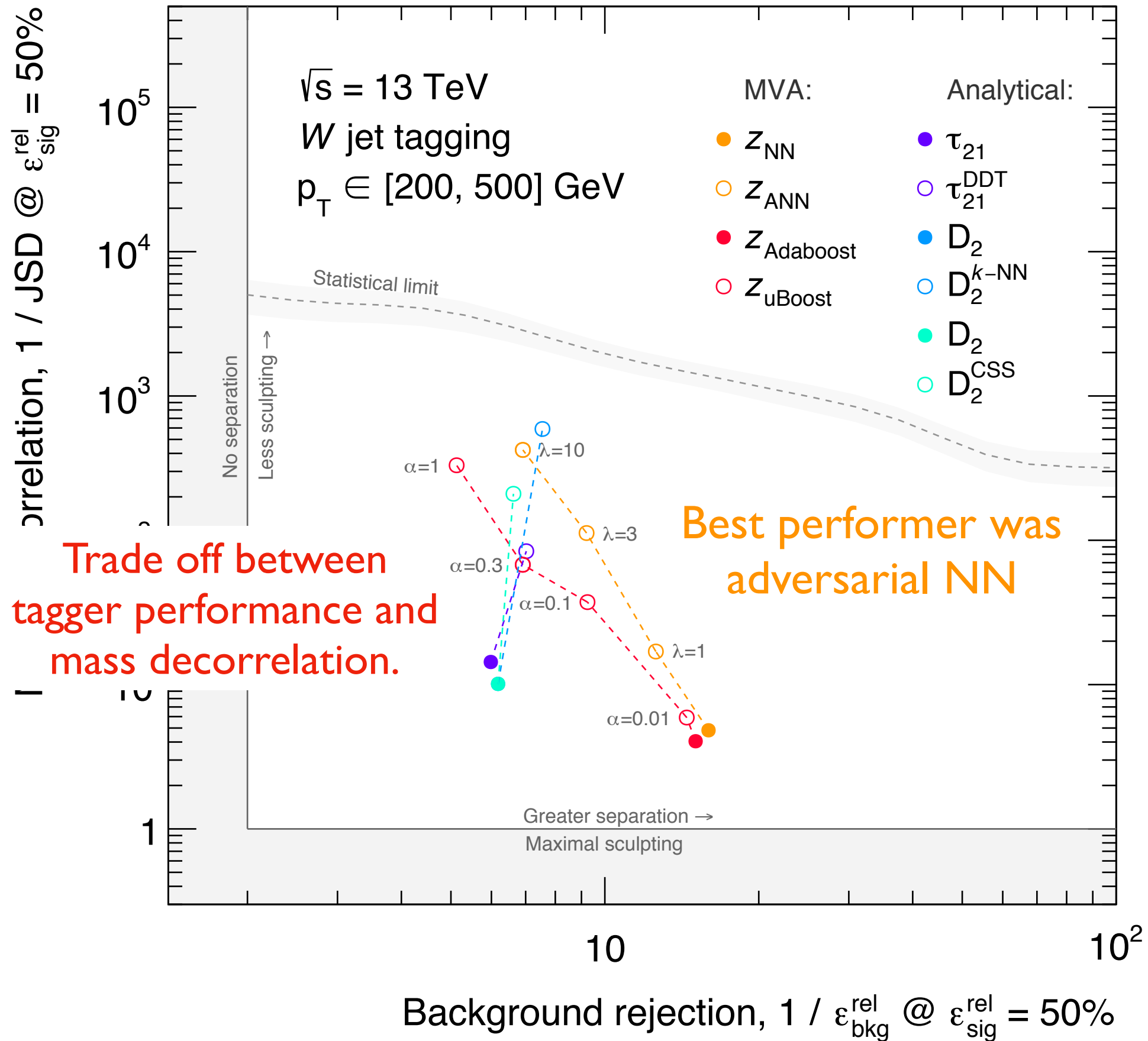
ATLAS Simulation Preliminary



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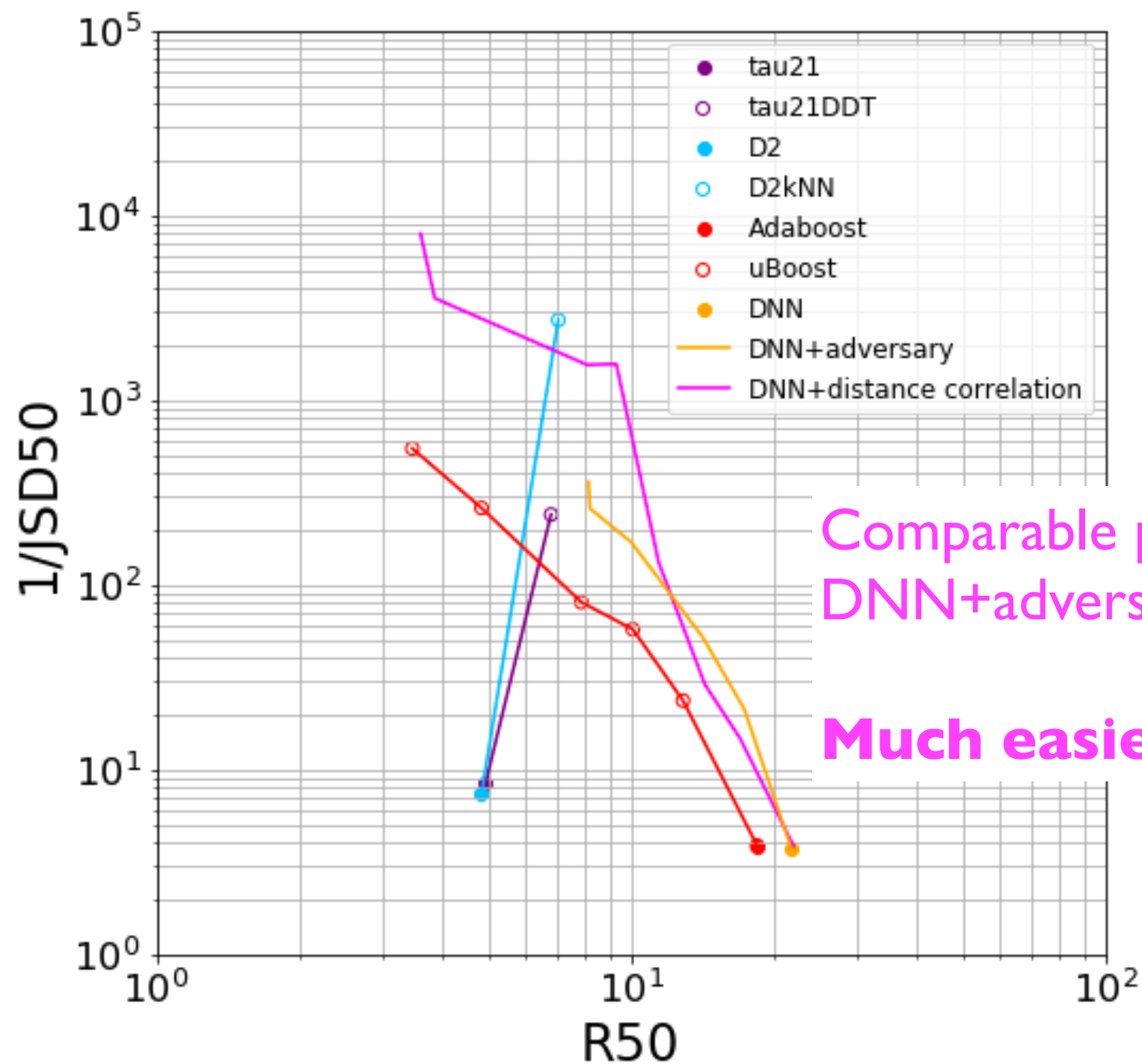


ATLAS Simulation Preliminary



Distance correlation (“DisCo”)

Work in progress with Gregor Kasieczka



Comparable performance to DNN+adversary.

Much easier to train.

Summary/Outlook

There is increasingly strong motivation to perform model-independent searches at the LHC. We need to ensure that we have not missed anything in the data.

However, attempts at model-independent searches have suffered from major drawbacks: an enormous trials factor, simplistic signal/background discrimination, and an over-reliance on simulation. Together, they have given the philosophy of model-independent searches a bad name.

In this talk we have explored new ideas for model-independent searches inspired by recent breakthroughs in unsupervised deep learning.

We have seen that:

- Deep autoencoders can find subtle signals in the data in a model agnostic way.
- Decorrelation is essential in combining data-driven background estimation with deep learning.
- A new method “DisCo” can more easily achieve state-of-the-art decorrelation in NN training.

Outlook

I think there is an exciting future for model-independent searches @ LHC!

- Comparison of different approaches to model-independent searches (AE, CWoLa, ... — work in progress with Pablo Martin & Ben Nachman)
- Careful study of the LEE in these different approaches
- More applications of DisCo
 - e.g. using it to improve the ABCD method (“Double DisCo” — work in progress with Gregor Kasieczka, Ben Nachman & Matt Schwartz)
- Real life applications of decorrelation?
 - E.g. making less discriminatory AIs for hiring/admissions/bail decisions/sentencing/... ?
- **New ideas for model-independent searches**
 - e.g. using recent breakthroughs in density estimation to search for anomalies
 - many more...?!

Overview

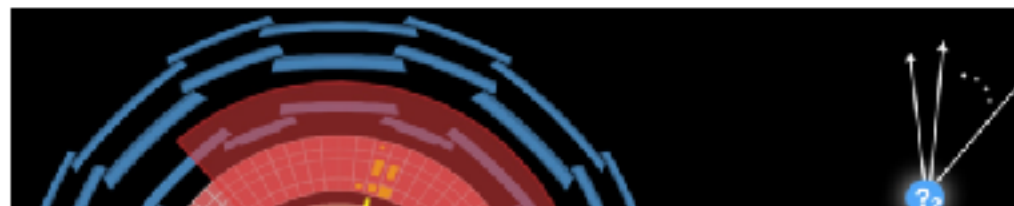
Timetable

Participant List

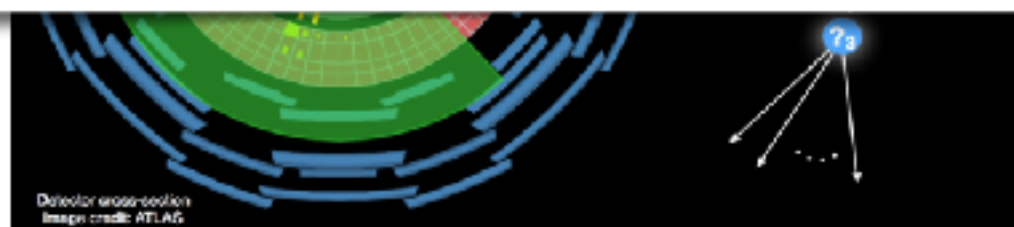
LHCOlympics2020

Slack channel

LHCOlympics2020



Come join us for the LHC Olympics 2020!



Despite an impressive and extensive effort by the LHC collaborations, there is currently no convincing evidence for new particles produced in high-energy collisions. At the same time, there has been a growing interest in machine learning techniques to enhance potential signals using all of the available information.

In the spirit of the first LHC Olympics (circa 2005-2006) [1st, 2nd, 3rd, 4th], we are organizing the 2020 LHC Olympics. Our goal is to ensure that the LHC search program is sufficiently well-rounded to capture "all" rare and complex signals. The final state for this olympics will be focused (generic dijet events) but the observable phase space and potential BSM parameter space(s) are large: all hadrons in the event can be used for learning (be it "cuts", supervised machine learning, or unsupervised machine learning).

For setting up, developing, and validating your methods, we provide background events and a benchmark signal model. You can download these from [this page](#). To help get you started, we have also prepared [simple python scripts](#) to read in the data and do some basic processing.

The final test will happen 2 weeks before the ML4Jets2020 workshop. We will release a new dataset where the "background" will be similar to but not identical to the one in the development set (as is true in real data!). The goal of the challenge is to see who can "best" identify BSM (yes/no, what mass, what cross-section) in the dataset. There are many ways to quantify "best" and we will use all of the submissions to explore the pros/cons of the various approaches.

Thanks for your attention!