Chaos and symmetry breaking in low-dimensional AdS/CFT

Bay Area Particle Theory Seminar - March 2017

Kristan Jensen (SFSU)

based on:

- arXiv:1605.06098 (PRL 117 (2016), 11, 116601)



### An old hope

Question: what is the minimal example of AdS/CFT?

Equivalently: what is the simplest consistent theory of quantum gravity on AdS?

Natural candidate: 3d pure gravity with negative cc

$$S_{3d} = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left( R + \frac{2}{L^2} \right)$$

Classical 3d gravity is simple:

Natural candidate: 3d pure gravity with negative cc

Classical 3d gravity is simple:

- 1. All solutions classified: boundary gravitons plus BTZ
- 2. May be recast as  $SO(2,1) \times SO(2,1)$ Chern-Simons theory

$$c = \frac{3L}{2G} = 24k$$

$$S_{3d} = \frac{k}{4\pi} \int \operatorname{tr}\left(A \wedge dA + \frac{2}{3}A \wedge A \wedge A\right)$$

Natural candidate: 3d pure gravity with negative cc

$$S_{3d} = \frac{k}{4\pi} \int \operatorname{tr}\left(A \wedge dA + \frac{2}{3}A \wedge A \wedge A\right) \qquad c = \frac{3L}{2G} = 24k$$

Witten 2007: Conjecture that for k=1,2,.., dual to "extremal CFT" with torus  $Z=|\chi_k(q)|^2$ 

$$\chi_k(q) = q^{-k} \prod_{n=2}^{\infty} (1 - q^n)^{-1} + O(q)$$

Natural candidate: 3d pure gravity with negative cc

$$Z = |\chi_k(q)|^2 \qquad c = \frac{3L}{2G} = 24k$$

k=1:  $\chi_1(q) = J(q)$  Modular J-function

this Z corresponds to known CFT! [Frenkel, Lepowsky, Meurman] So-called Monster CFT

Natural candidate: 3d pure gravity with negative cc

$$Z = |\chi_k(q)|^2 \qquad c = \frac{3L}{2G} = 24k$$

k=1: 
$$\chi_1(q) = J(q)$$
 Modular J-function

Lightest states: 196883 operators of dimension 2

Identifying as BH creation ops, get

$$S = \ln 196883 \approx 12.19$$
 vs.  $S_{BH}(k=1) = 4\pi \approx 12.57$ 

Natural candidate: 3d pure gravity with negative cc

$$Z = |\chi_k(q)|^2 \qquad c = \frac{3L}{2G} = 24k$$

k=1:  $\chi_1(q) = J(q)$  Modular J-function

k>1: not known if corresponding CFT exists, some evidence against

### What about AdS<sub>2</sub>/CFT<sub>1</sub>?

Simplest possible setting for the duality

And ubiquitous: AdS<sub>2</sub> generically appears as near-horizon of SUSY black holes

### What about AdS<sub>2</sub>/CFT<sub>1</sub>?

Simplest possible setting for the duality

And ubiquitous: AdS<sub>2</sub> generically appears as near-horizon of SUSY black holes

PROBLEM: Neither AdS<sub>2</sub> gravity nor CFT<sub>1</sub> exist

### The problem

AdS<sub>2</sub>: cannot support finite-energy excitations,  $\langle E \rangle = 0$ 

CFT<sub>1</sub>: One runs into a paradox of [Polchinski]

scale-invariant density of states in 1d is

$$\rho(E) = e^{S_1} \delta(E) + \frac{e^{S_2}}{E}$$

### The problem

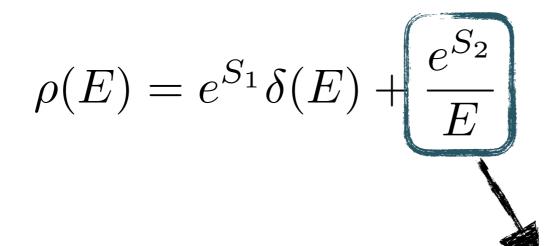
scale-invariant density of states in 1d is

$$\rho(E) = e^{S_1} \delta(E) + \frac{e^{S_2}}{E}$$

Zero-energy states
No dynamics!

### The problem

scale-invariant density of states in 1d is



Divergent at E=0

Z does not exist!

#### Goal for this talk:

Obtain sensible notion of AdS<sub>2</sub>/CFT<sub>1</sub>

### Outline

- 1. Motivation
- 2. NAdS<sub>2</sub>/NCFT<sub>1</sub>
- 3. Chaos
- 4. Parting words

### Outline

1. Motivation

2. NAdS<sub>2</sub>/NCFT<sub>1</sub>

3. Chaos

4. Parting words

### Dilaton gravity

Pure 2d gravity is topological 
$$\int d^2 x \sqrt{-g} \, R = 4\pi \chi$$

Reduction to 2d leads to dilaton gravity

$$S = \frac{1}{2\kappa^2} \int d^2x \sqrt{-g} \left(\varphi R + U[\varphi]\right) + S_{\text{matter}}$$

### Dilaton gravity

Reduction to 2d leads to dilaton gravity

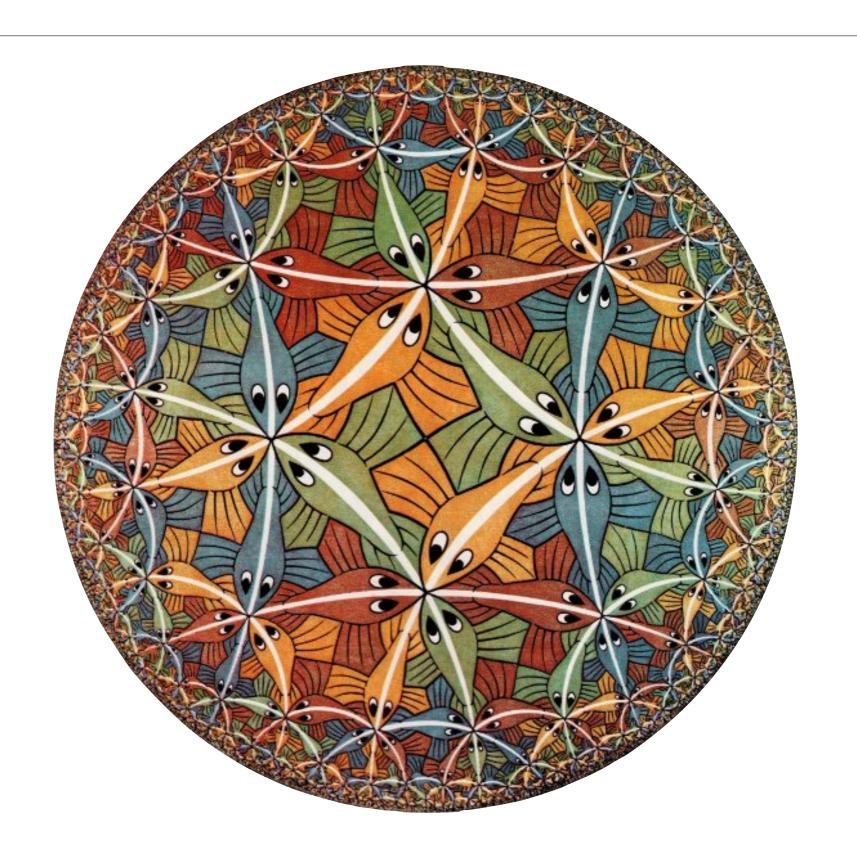
$$S = \frac{1}{2\kappa^2} \int d^2x \sqrt{-g} \left(\varphi R + U[\varphi]\right) + S_{\text{matter}}$$

General two-derivative theory characterized by U

AdS<sub>2</sub> solutions with constant dilaton at roots of U

$$U[\varphi_0] = 0, \qquad U'[\varphi_0] = \frac{2}{L^2}$$

# $AdS_2$



Most general AdS<sub>2</sub> spacetime characterized by a free function of boundary time

$$g = -r^2 \left(1 + \frac{h(t)}{r^2}\right)^2 dt^2 + \frac{dr^2}{r^2}$$
  $\varphi = \varphi_0$ 

Most general AdS<sub>2</sub> spacetime characterized by a free function of boundary time

Convenient to redefine 
$$h(t) = \frac{1}{2} \{f(t), t\}$$

Most general AdS<sub>2</sub> spacetime characterized by a free function of boundary time

Convenient to redefine 
$$h(t) = \frac{1}{2}\{f(t), t\}$$

- 1. f(t) acts as conformal transformation on bdy
- 2. For any f(t), invariant under  $PSL(2,\mathbb{R})$

$$f(t) o rac{af(t) + b}{cf(\tau) + d}$$

Convenient to redefine 
$$h(t) = \frac{1}{2} \{f(t), t\}$$

- 1. f(t) acts as conformal transformation on bdy
- 2. For any f(t), invariant under  $PSL(2,\mathbb{R})$

boundary conformal symmetry

= Virasoro symmetry with c=0

### NAdS<sub>2</sub> spacetimes

Name due to [Maldacena, Stanford]

$$U[\varphi] \approx 2(\varphi - \varphi_0)$$
 also admits linear dilaton solutions

### NAdS<sub>2</sub> spacetimes

Name due to [Maldacena, Stanford]

Arises as IR endpoint of ANY holographic RG flow ending in AdS<sub>2</sub>

### NAdS<sub>2</sub> spacetimes

Name due to [Maldacena, Stanford]

$$U[\varphi] \approx 2(\varphi - \varphi_0)$$

$$g = -r^2 dt^2 + \frac{dr^2}{r^2}$$

$$\varphi = \ell r + \varphi_0$$

Moduli f(t) become pseudo-moduli: Virasoro  $\longrightarrow PSL(2,\mathbb{R})$ 

Dual to nearly conformal QM, dubbed NCFT<sub>1</sub>

# Fluid/gravity

Write NAdS<sub>2</sub> in infalling Eddington-Finkelstein coordinates

$$g = -(r^2 + 2\{f(t), t\}) dt^2 + 2dtdr + O(\varepsilon)$$
$$\varphi = \varphi_0 + \varepsilon \ell r + O(\varepsilon^2) \qquad T_{\mu\nu} = O(\varepsilon)$$

Solve bulk eoms exactly near AdS<sub>2</sub>; can rewrite as equation on the boundary:

$$\dot{E}=\dot{\lambda}\langle O
angle$$
 with  $\left[E=-rac{\ell}{\kappa^2}\{f(t),t\}+(1-\Delta)\lambda\langle O
angle
ight]$ 

#### Schwarzian action

Both follow from action where f(t) is fundamental field:

$$S = -\frac{\ell}{\kappa^2} \int dt \left\{ f(t), t \right\} + W[\lambda(t); f(t)]$$

with 
$$\{f(t),t\} = \frac{f'''(t)}{f'(t)} - \frac{3}{2} \left(\frac{f''(t)}{f'(t)}\right)^2 =$$
Schwarzian derivative

[KJ], [Maldacena, Stanford, Yang], [Engelsoy, Martens, Verlinde]

also obtained in the low-energy limit of Sachdev-Ye-Kitaev models [Maldacena, Stanford]

#### Schwarzian action

Both follow from action where f(t) is fundamental field:

$$S = -\frac{\ell}{\kappa^2} \int dt \{f(t), t\} + W[\lambda(t); f(t)] + O(\ell \lambda^2, \ell^2)$$

Leading interactions consistent with

Virasoro 
$$\longrightarrow PSL(2,\mathbb{R})$$

# The good and the ugly

(Very) good:

Rewrite two-derivative NAdS<sub>2</sub> gravity as 1d theory

# The good and the ugly

(Very) good:

Rewrite two-derivative NAdS<sub>2</sub> gravity as 1d theory

Ugly:

Resulting Schwarzian theory is sick in isolation [Stanford, Witten]

(situation recalls [Maloney, Witten] prescription for partition function of pure 3d gravity)

### Outline

- 1. Motivation
- 2. NAdS<sub>2</sub>/NCFT<sub>1</sub>
- 3. Chaos
- 4. Parting words

### Lyapunov exponent

Quantum mechanical analogue of Lyapunov exponent, characterizing early-time chaotic growth:

$$\langle [W(t), V(0)]^2 \rangle_{\beta} \sim e^{\lambda_L t}$$

### Lyapunov exponent

Quantum mechanical analogue of Lyapunov exponent, characterizing early-time chaotic growth:

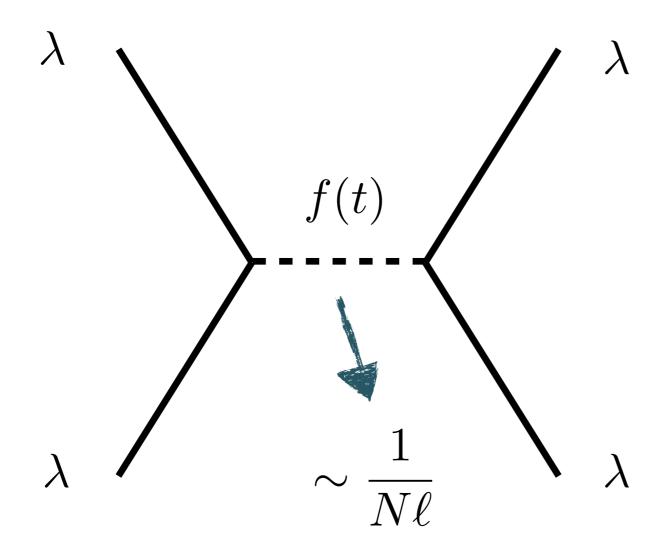
$$\langle [W(t), V(0)]^2 \rangle_{\beta} \sim e^{\lambda_L t}$$

#### Two properties:

- I. [Maldacena, Shenker, Stanford]  $\lambda_L \leq rac{2\pi}{eta}$
- 2. Dual to Einstein gravity is maximally chaotic [Shenker, Stanford]

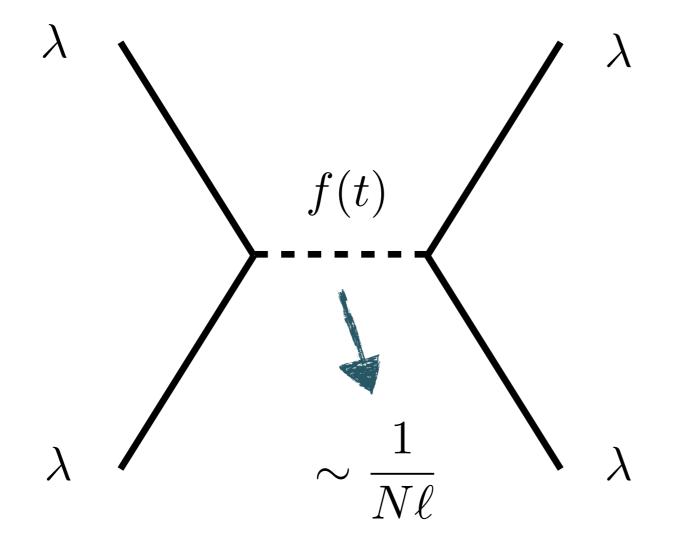
# Lyapunov exponent in NAdS<sub>2</sub>/NCFT<sub>1</sub>

Compute Euclidean four-point function at tree-level:



### Lyapunov exponent in NAdS<sub>2</sub>/NCFT<sub>1</sub>

Compute Euclidean four-point function at tree-level:



Analytically continue to get out-of-time-ordered four-point function

$$\lambda_L = \frac{2\pi}{eta}$$

### Outline

- 1. Motivation
- 2. NAdS<sub>2</sub>/NCFT<sub>1</sub>
- 3. Chaos
- 4. Parting words

# A sigma model for AdS<sub>3</sub> gravity

WIP with Jordan Cotler (Stanford)

It is possible to rewrite pure classical AdS<sub>3</sub> gravity as an unconventional sigma model from  $\mathcal{M}_2 \to \mathcal{M}_2$ 

# A sigma model for AdS3 gravity

WIP with Jordan Cotler (Stanford)

The action may be regarded as an action for an "exact" hydrodynamics whose gradient expansion truncates at second order in derivatives

$$S = \frac{c}{24\pi} \int d^2 \sigma \sqrt{-g} \left( 4\pi^2 T^2 - \frac{(\partial T)^2}{T^2} - \ln T R \right)$$

### A sigma model for AdS3 gravity

WIP with Jordan Cotler (Stanford)

$$S = \frac{c}{24\pi} \int d^2 \sigma \sqrt{-g} \left( 4\pi^2 T^2 - \frac{(\partial T)^2}{T^2} - \ln T R \right)$$

Torus partition function?

Coupling to matter and Virasoro blocks?

Stay tuned!

Thank you!